

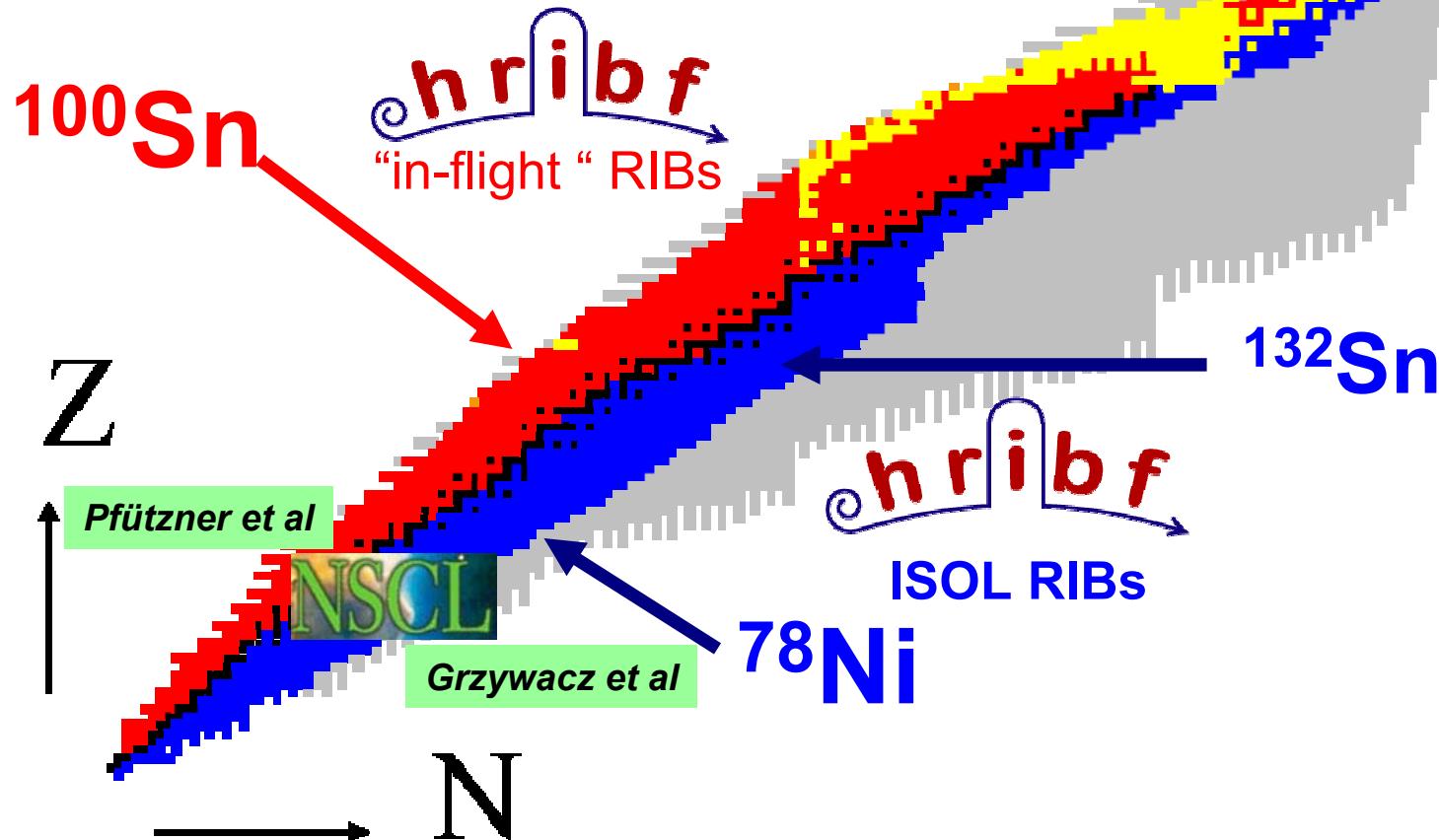
# *Perspektywy badań rozpadów egzotycznych jąder w laboratorium hribf (Oak Ridge, TN, USA)*

**K. P. Rykaczewski**

*perspektywy badań z dalszym udziałem polskich współpracowników*

Warszawa, Kraków, Łódź ...

hribf = *Holifield Radioactive Ion Beam Facility*





*HRIBF users are getting their decay spectroscopy projects done !*

- M. Karny (Warszawa) + 26, "proton emitters  $^{141m,gs}\text{Ho}$ ", submitted to Phys. Lett. B (**+ habilitacja ?**)  
A. Korgul (Warszawa) + 14, "Toward  $^{100}\text{Sn}$ ", Phys. Rev. C77, 034301, 2008  
J. Winger (Mississippi) + 17, "Failure of  $^{78}\text{Ni}$  core?", World Scientific, in print  
S. Ilyushkin (Mississippi) + 17, " $\beta$  and  $\beta\text{n}$ -decays  $^{76-79}\text{Cu}$ ", World Scientific, in print, (**+PhD**)  
J. Winger .. A. Korgul, (Mississippi-Warszawa) + 16 coauthors, Acta. Phys. Pol.B39, 525, 2008  
S.N.Liddick (UTK) + 22 coauthors, EPJ Special Topics 150, 131, 2007  
C. Mazzocchi (UTK-Milano) + 16 coauthors, Phys. Rev. Lett. 98, 212501, 2007   
S.N. Liddick (UNIRIB) + 22 coauthors, Phys. Rev. Lett., 97, 082501, 2006  
M.N.Tantawy (UT) + 23 coauthors, PR C73, 024316, 2006 (**+PhD**)  
J.C. Batchelder (UNIRIB) + 21 coauthors, PR C72, 044306, 2005  
R.K.Grzywacz (UTK/ORNL) + 12 coauthors, EPJ A25, s01, 145, 2005  
T.N.Ginter (Vanderbilt/MSU) +12 coauthors, PR C68, 034330, 2003 (**+PhD**)  
M.Karny (Warszawa) + 17 coauthors, Phys.Rev.Lett. 90, 012502, 2003  
W.Królas (Kraków) + 24 coauthors, PR C65, 031303(R), 2002  
J.J. Ressler (Maryland) + 17 coauthors, Phys. Rev. Lett. 84, 2104, 2000 (**+PhD**)  
A.Piechaczek (LSU) + 10 coauthors Phys. Rev. C 61, 047306, 2000

*Local decay spectroscopy team : 1.25 FTE (only !)*

# Diversity ?? No problem !



Oak Ridge, Walentynki 2008  
badanie rozpadu  $^{112}\text{Cs}$



**The shell structure description of stable nuclei may not apply as readily to nuclei outside the valley of stability !**

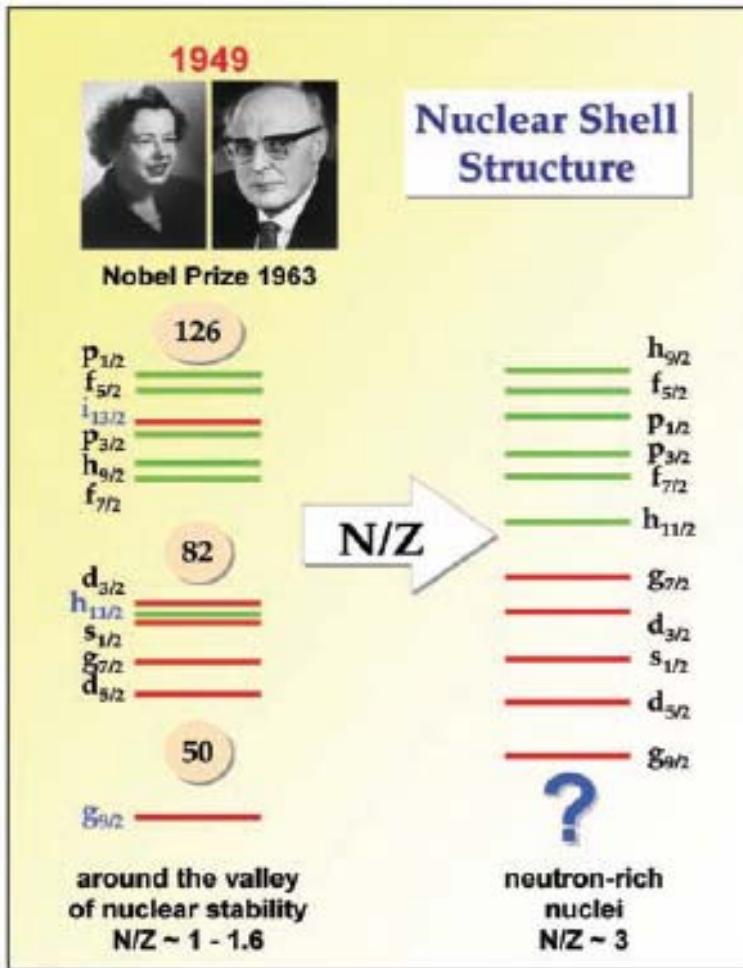
## “Shell Structure of Exotic Nuclei”

J.Dobaczewski, N.Michel, W.Nazarewicz, M.Płoszajczak, J.Rotureau  
Progr. Part.Nucl. Phys. 59, 432, 2007

M. Goeppert-Mayer  
J. Hans D. Jensen

**M.Goeppert-Mayer**  
*Phys.Rev. 75, 1969,*  
**1949**

**Haxel, Jensen, Süss**  
*Phys. Rev. 75, 1766,*  
**1949**



**stable Ni**  
N/Z~1.1

**78Ni**    “**112Ni , 200Sn** “  
N/Z=1.8



J.Dobaczewski, I.Hamamoto,  
W.Nazarewicz, J.A.Sheikh  
*Phys. Rev. Lett. 72, 981, 1994*

J.Dobaczewski, W Nazarewicz,  
et al.,  
*Phys. Rev. C 53, 2809, 1996*

**“Scientific Opportunities  
with  
a Rare-Isotope Facility  
in the United States”**

[www.national-academies.org](http://www.national-academies.org)  
[www.nap.edu](http://www.nap.edu)  
2007

**HRIBF (Oak Ridge) :**  
staramy się zmierzyć i zrozumieć  
strukturę jąder najbardziej odległych od ścieżki stabilności beta  
(Lysekil , Szwecja, 1966)

Odkrywamy i badamy nowe nuklidy (2000+):  
nowe radioaktywnosci protone (6) i emitery alfa powyżej  $^{100}\text{Sn}$  (3),  
jak i nowe jądra neutrono-nadmiarowe takie jak  $^{79}\text{Cu}$  i  $^{85}\text{Ga}$   
(zaakceptowane projekty : rozпадy nowych jąder  $^{80-82}\text{Cu}$ ,  $^{86,87}\text{Ge}$  ,  $^{86,87}\text{Ga}$ ,  $^{88}\text{As}$ )

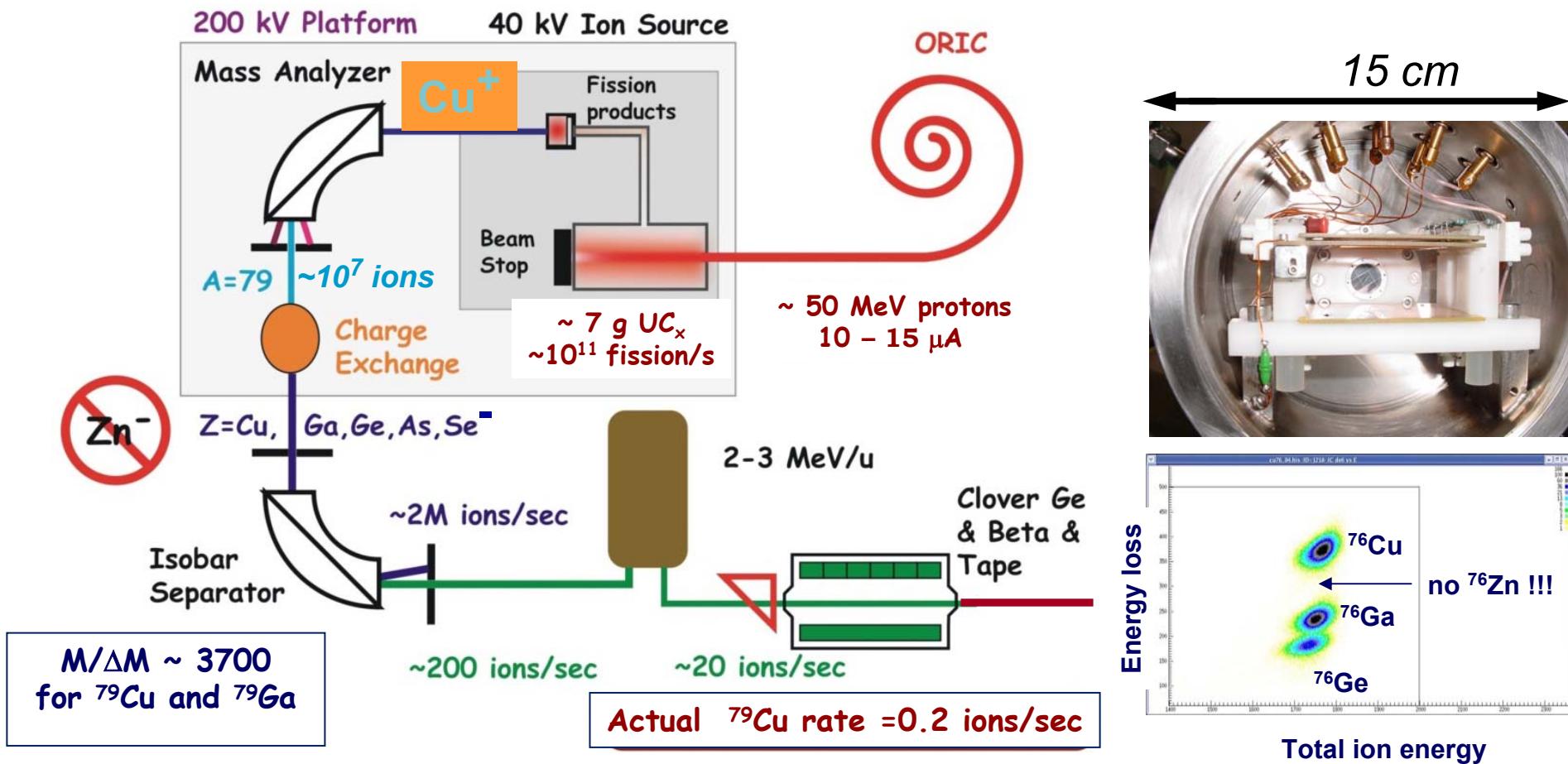
Staramy się znaleźć i zrozumieć :  
efekty związane z dużą asymetrią liczb protonów  $Z$  i neutronów  $N$ ,  
wpływ asymetrii  $N/Z$  na przebieg poziomów jednocząstkowych  
neutrony  $1g_{7/2}$  vs  $2d_{5/2}$  ( $N=51$   $^{101}\text{Sn}^*$ ,  $N=53$   $^{105}\text{Te}^*$ ),  
protony  $1f_{7/2}$  [ $Z=28$ ]  $1f_{5/2}$  vs  $2p_{3/2}, 2p_{1/2}$  ( $Z=29$  Cu)  
[ $N=50$ ] neutrony  $2d_{5/2}$  vs  $3s_{1/2} \rightarrow [N=58]$   
*struktura stanów identyfikowanych jako “jednocząstkowe” (! ?)*

### “zastosowania”:

Astrofizyka :  $r$ -proces powyżej  $^{78}\text{Ni}$  oraz  $rp$ -proces powyżej  $^{100}\text{Sn}$   
Dane o emisji opóźnionych neutronów i  $T_{1/2}(\beta)$  dla nowych generacji reaktorów

# Neutron-rich beams at HRIBF (Oak Ridge)

## production, mass-separation and „ranging-out”



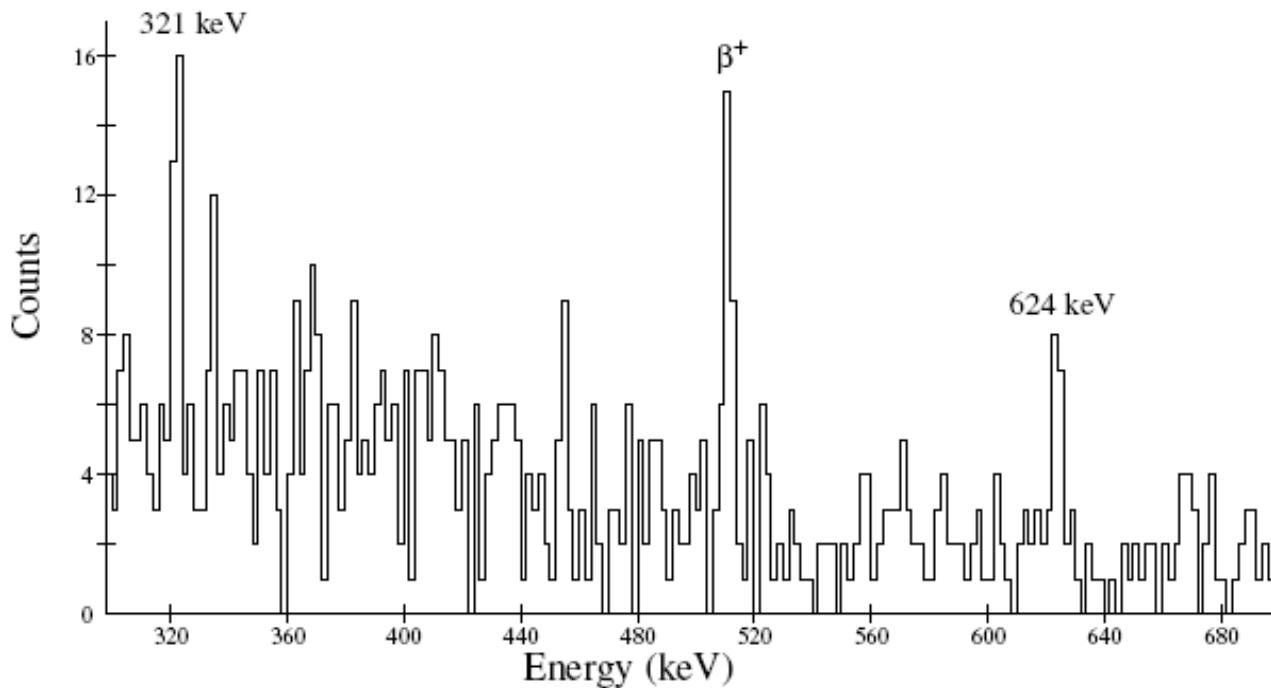
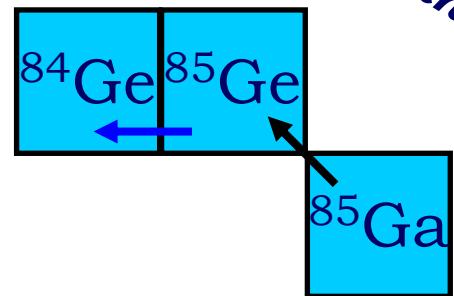
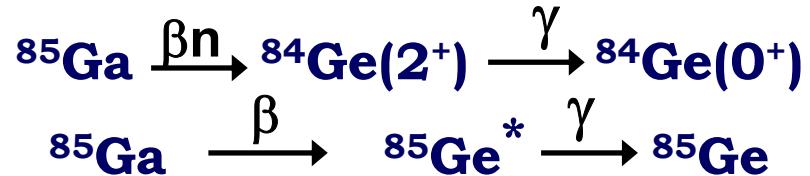
C.J.Gross et al., Eur.Phys.Jour. A25, s01, 115 (2005)

J.A.Winger et al., in contr. to Int. Nucl. Phys. Conf.(INPC) , Japan, June 2007

*preliminary*

$^{85}\text{Ga}$  decay,  $T_{1/2} = ???$  (< 100 ms)

rate ~ 0.1 pps, ~ 6 hours counting

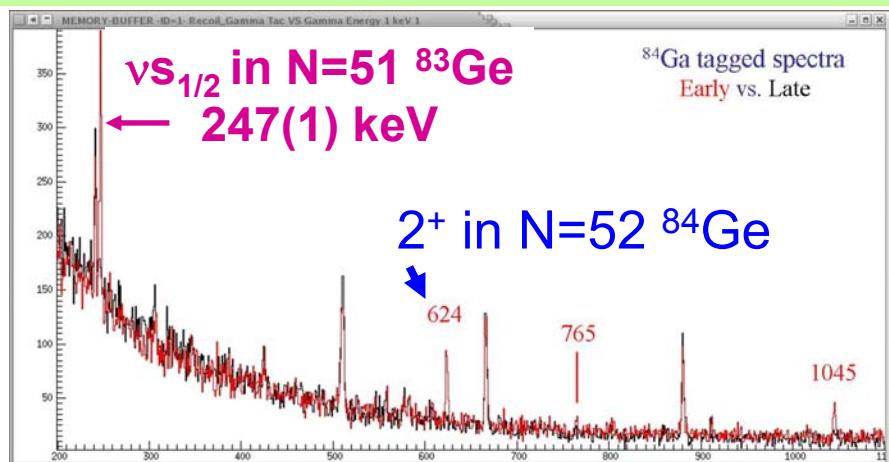
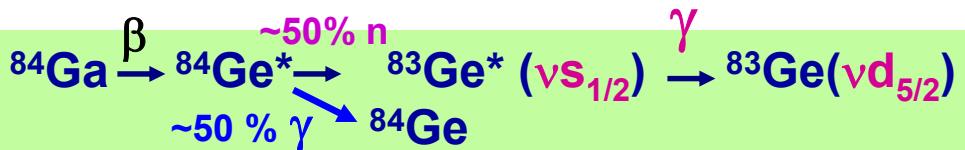


*Ion-correlated (0-200 ms)  $\beta$ -gated gamma energy spectrum (2 keV/ch)*

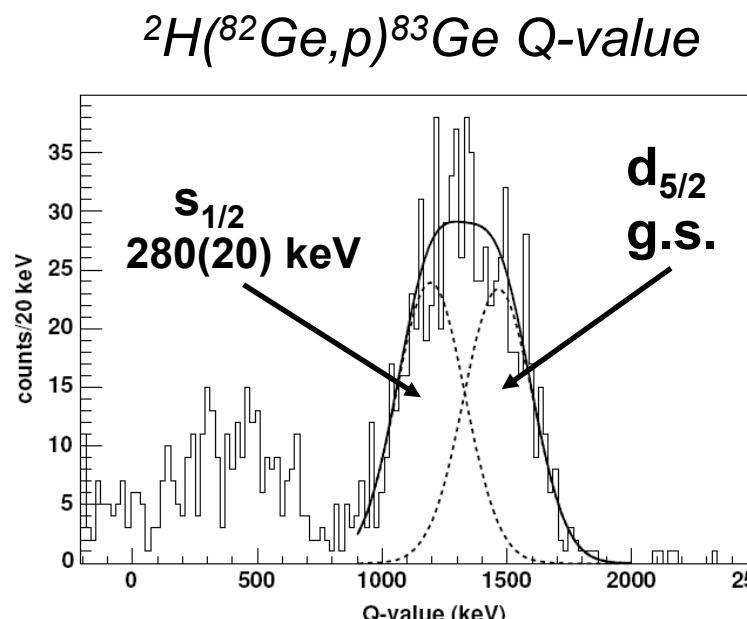
**J.Winger, S.Ilyushkin et al., HRIBF,**

*Int. Conf. on Fission and Properties of Neutron-rich nuclei”, Nov. '07, Sanibel Island, FL*

Nov'06 : 23 hours experiment with ~ 2 pps of N=53  $^{84}\text{Ga}$  ( $T_{1/2}=85\text{ ms}!!$ )



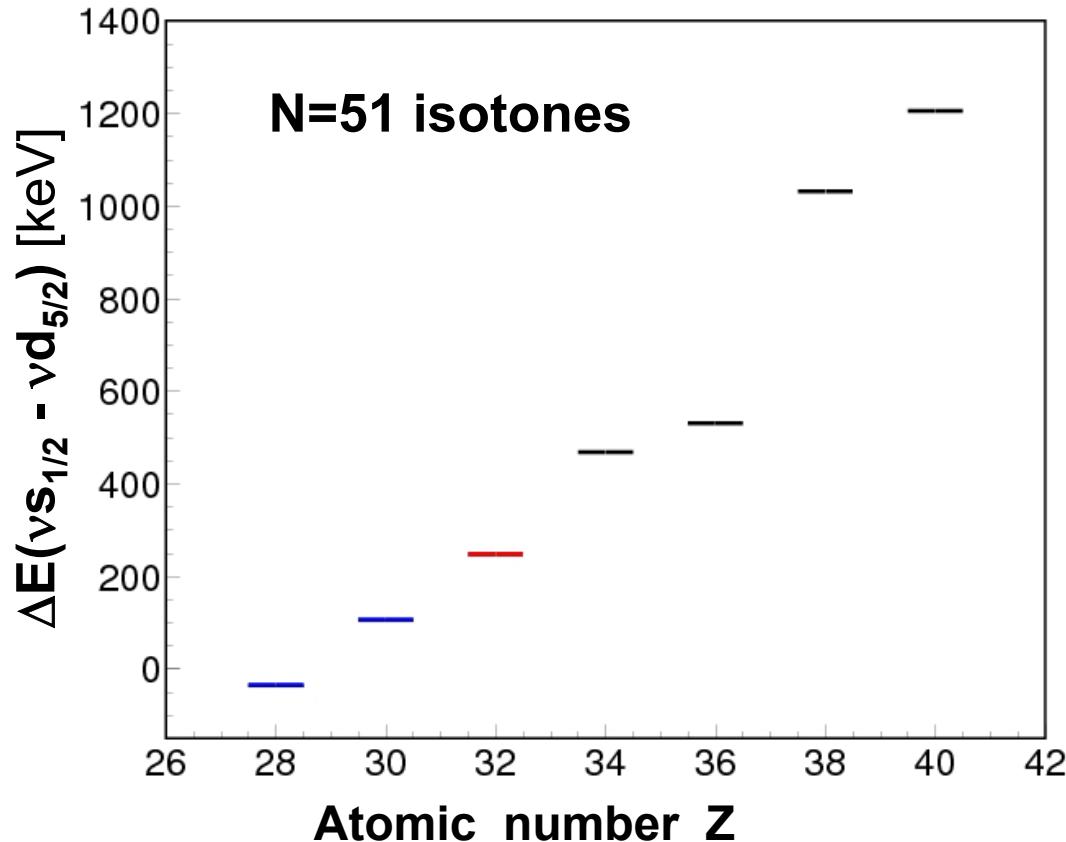
Ion-correlated  $\beta\gamma$ -spectrum (1 keV/ch)



Thomas et al., PR C71, 021302, 2005

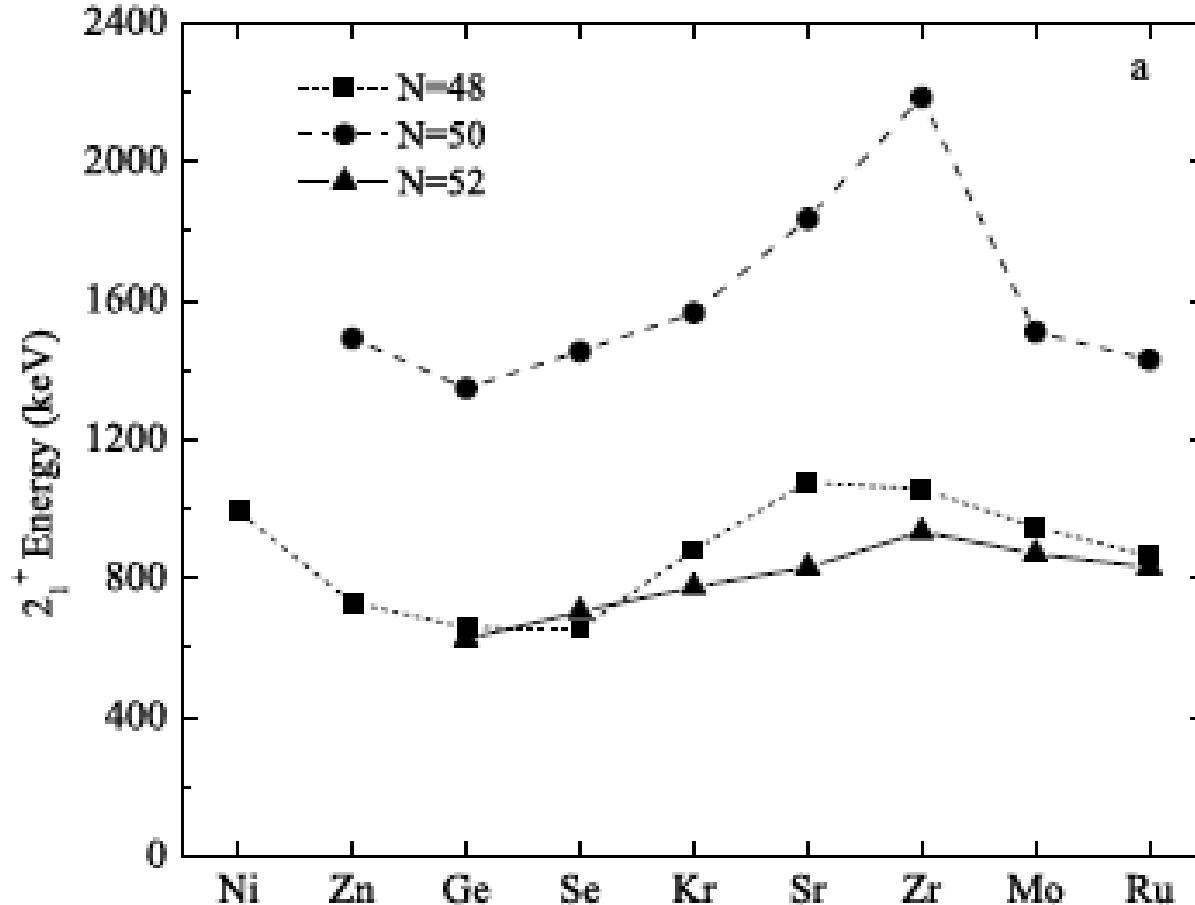
$\nu s_{1/2}$  and  $\nu d_{5/2}$  orbitals within ~200 keV  
new N=58 subshell closure in n-rich nuclei ??

**Our result :  $\nu s_{1/2}$  and  $\nu d_{5/2}$  orbitals within 247 keV in Z=32  $^{83}\text{Ge}$**   
**new N=58 subshell closure in n-rich nuclei ??**

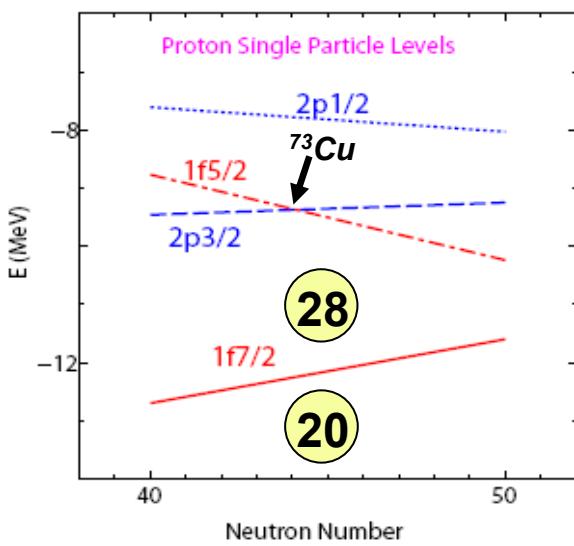
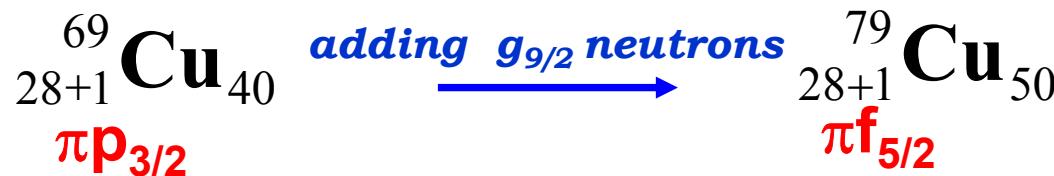


**extrapolation of the  $\nu s_{1/2}$ - $\nu d_{5/2}$  energies suggests  $\mu s - ms$  E2 isomers in  $^{81}\text{Zn}$  and  $^{79}\text{Ni}$**   
 **$\nu s_{1/2}$  isomeric halo in  $^{81}\text{Ni}$  ( $S_n \sim 170$  keV) ?**

*Zmierzone energie pierwszych stanów  $2^+$   
w parzysto-parzystych izotonach  $N=48$ ,  $N=50$  oraz  $N=52$*



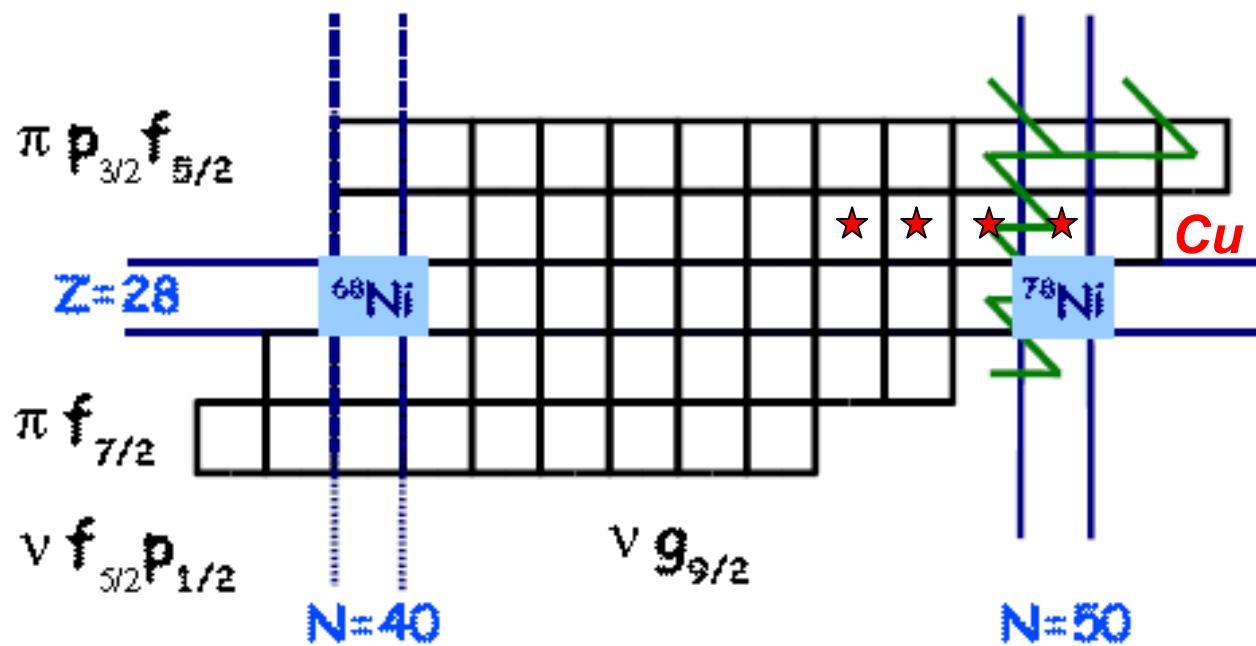
★ *HRIBF studies of  $^{76}\text{Cu}$  -  $^{79}\text{Cu}$*



*T. Otsuka et al.,*

PRL 95, 232502, 2005

*linear extrapolation  $f_{5/2}$  vs  $f_{7/2}$   
no Z=28 gap near  $^{83}\text{Cu}$   
and Z=34 gap appears !!!*



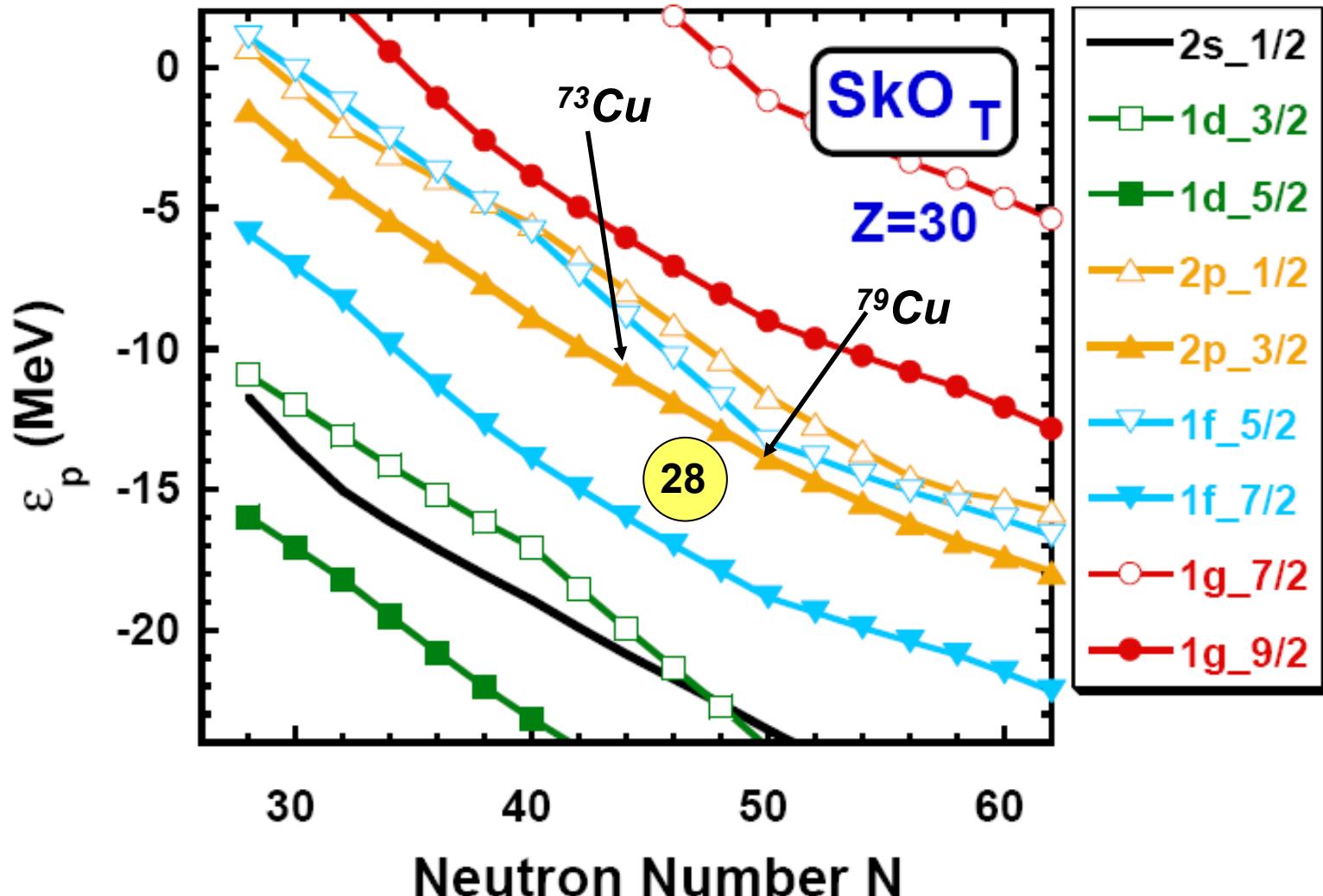
— “r-process path”

Jacek Dobaczewski, January 2008

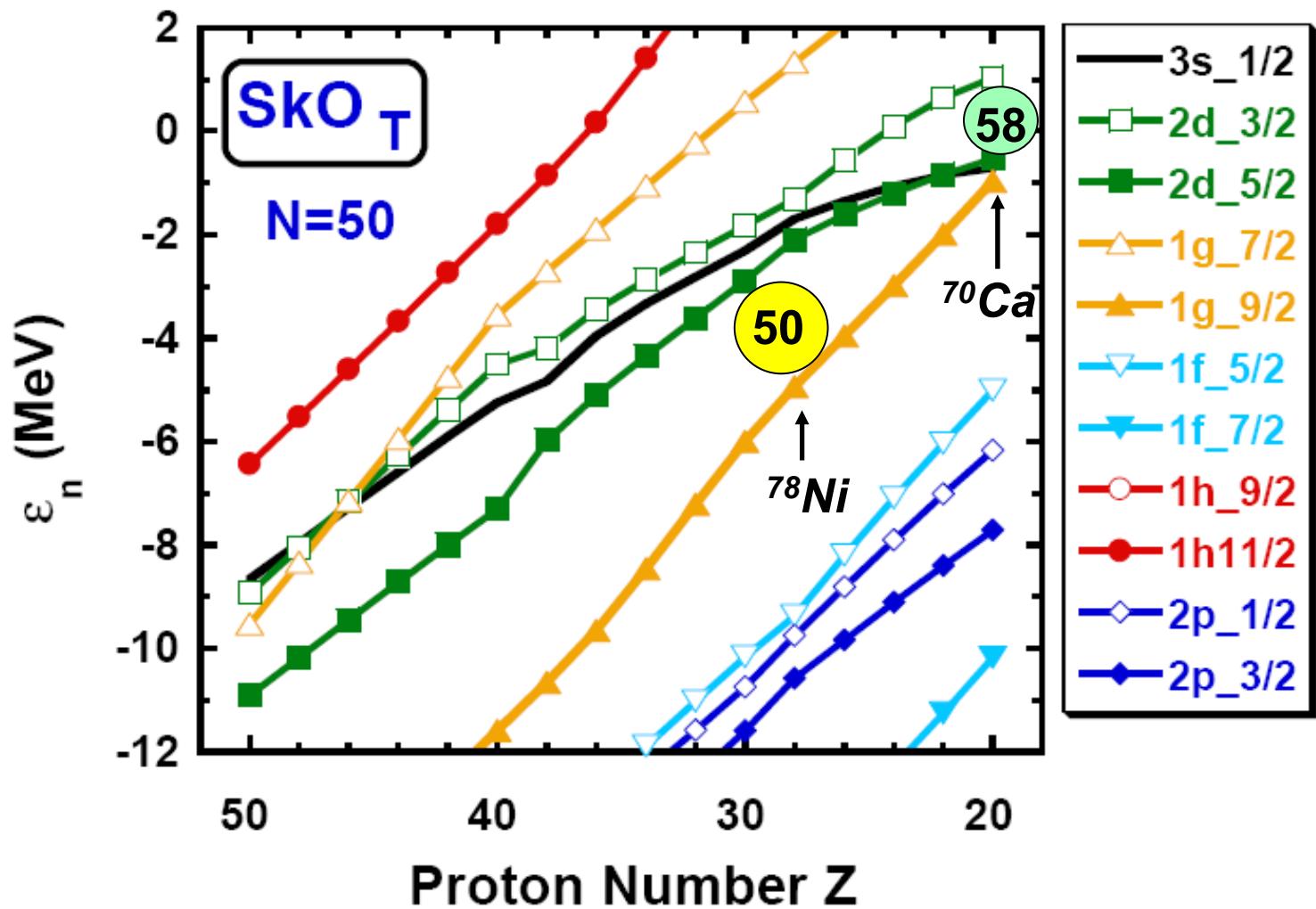
*"first attempt"* (to be revisited, refined, retuned, refitted, recalculated)

Physics :  $\pi 2p_{3/2}$  vs  $\pi 1f_{5/2}$  crossing (?) and an evolution of Z=28 gap

HRIBF contribution :  $\beta$ -decays above Z=28,  $\beta$ -NMR, 2+ energies ...



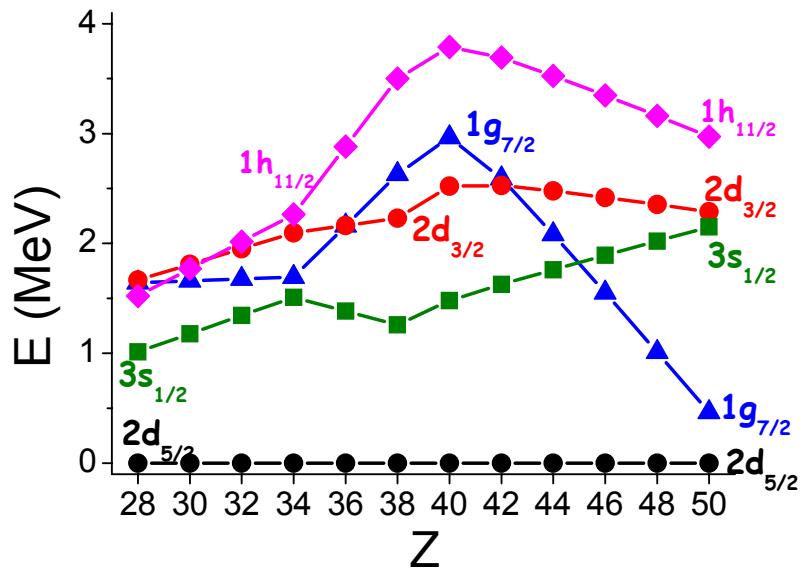
Jeff Winger et al.,  
 2007 Sanibel conference contr.: "Failure of the  $^{78}\text{Ni}$  core for  $Z>28$ ,  $N>50$  ?"  
 (low  $2+$  energies like 624 keV in  $Z=32$ ,  $N=52$   $^{84}\text{Ge}$ )



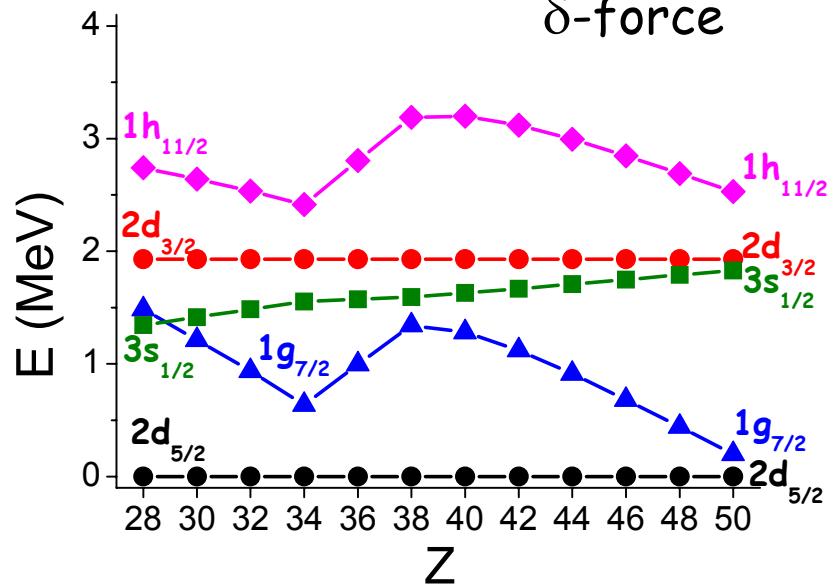
Jacek Dobaczewski, Jan 2008, to be revisited etc. ( $N=52 - 58$  calculations !)

**Note :** relative energy of  $\nu s_{1/2}$  vs  $\nu d_{5/2}$  state !!

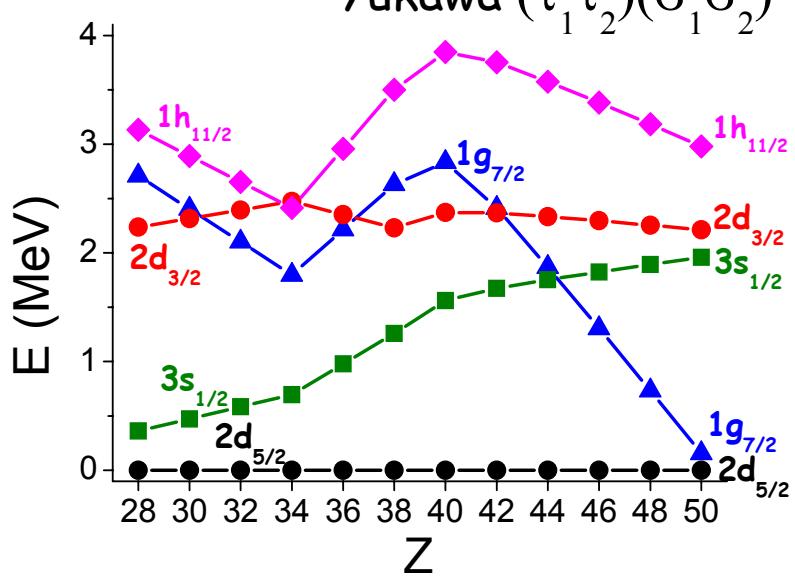
realistic interaction

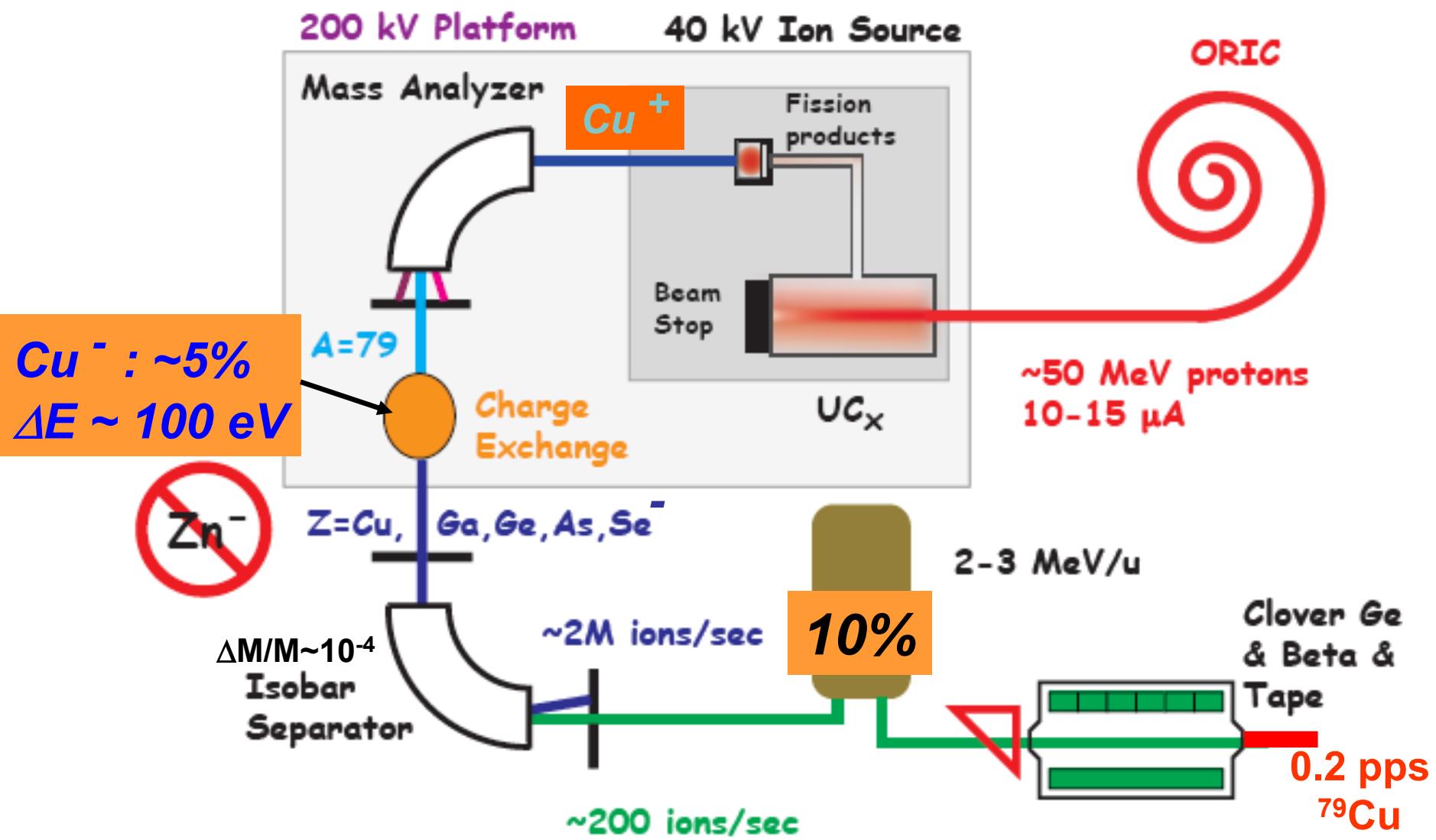


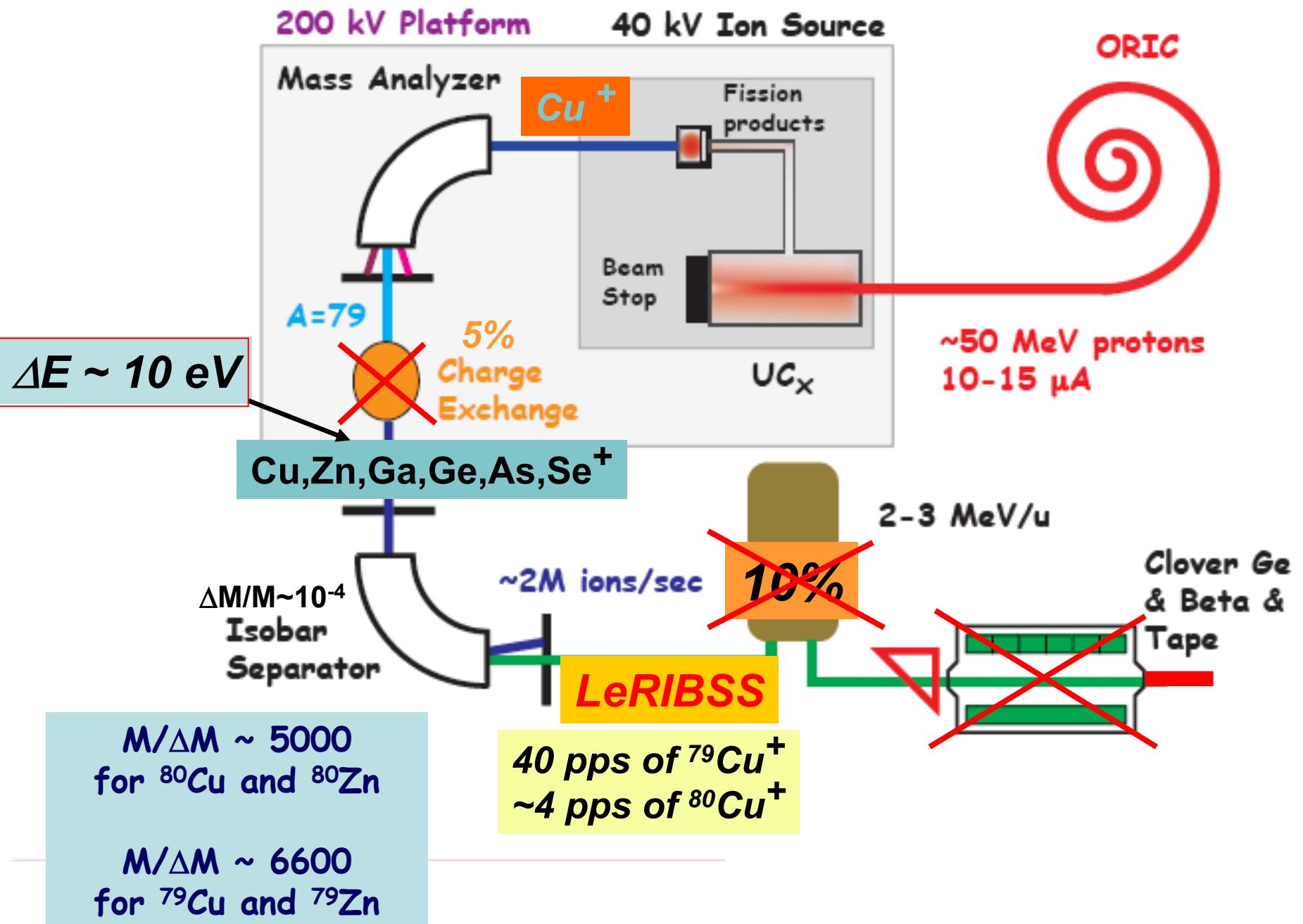
$\delta$ -force



Yukawa ( $\tau_1 \tau_2$ )( $\sigma_1 \sigma_2$ )



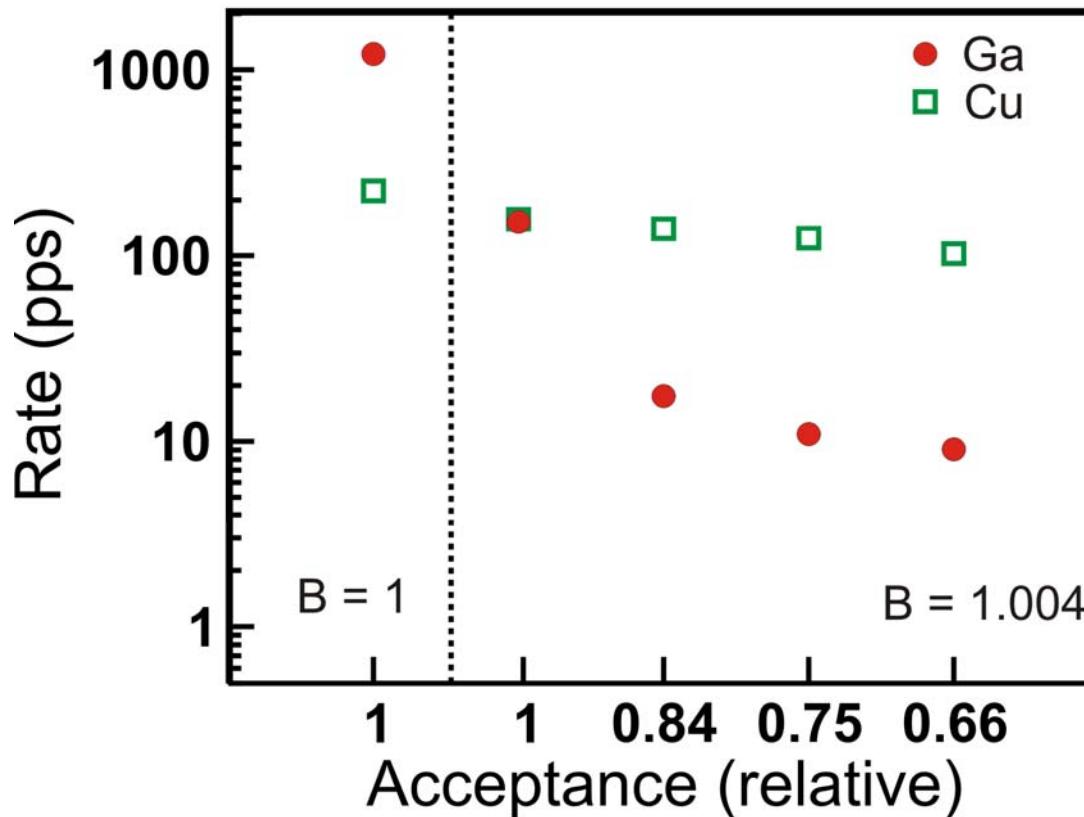




HRIBF high-resolution RIB injector magnet  $\vec{B}$   
(designed :  $\Delta M/M \sim 1 : 20\,000$ )

rate of  $A=76$  isobars  $\sim 10^5$  pps  $\longrightarrow$  “ $\vec{B}$ -optimized“ rate ( $^{76}\text{Cu}^-$ ) $\sim 220$  pps

Experimental rates at mass 76



for  $^{76}\text{Cu}$  -  $^{76}\text{Ga}$   $\Delta M/M \sim 1 : 4600$

# Low-energy Radioactive Ion Beam Spectroscopy Station LeRIBSS

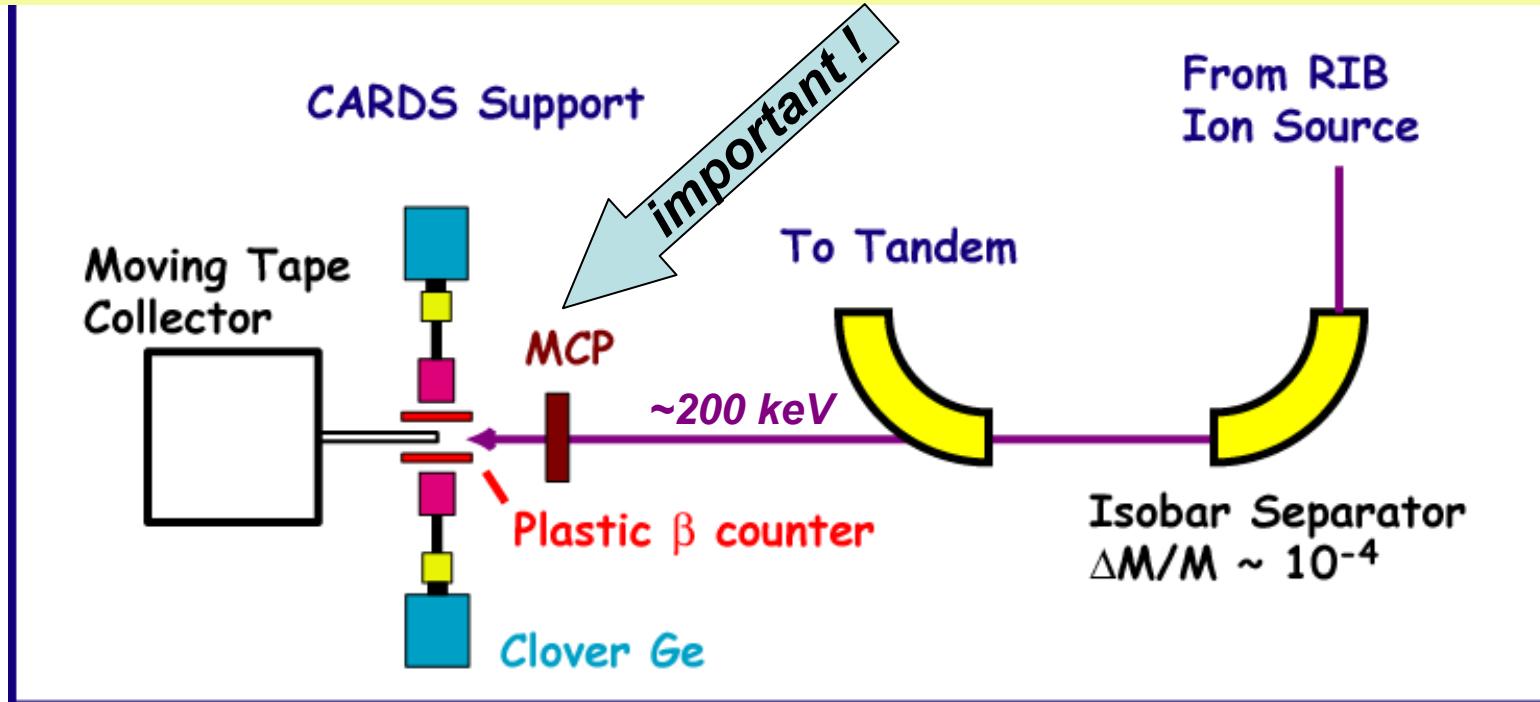
**Factor 20 to 1000 improvement in RIBs intensity**

**no Tandem – 10, no Charge Exchange – 2++ ( e.g. Cu,Ga –  $10 \times 20 = 200$ )**

**negative AND positive ~ 200 keV ions from IRIS-1 and 250 keV from IRIS-2**

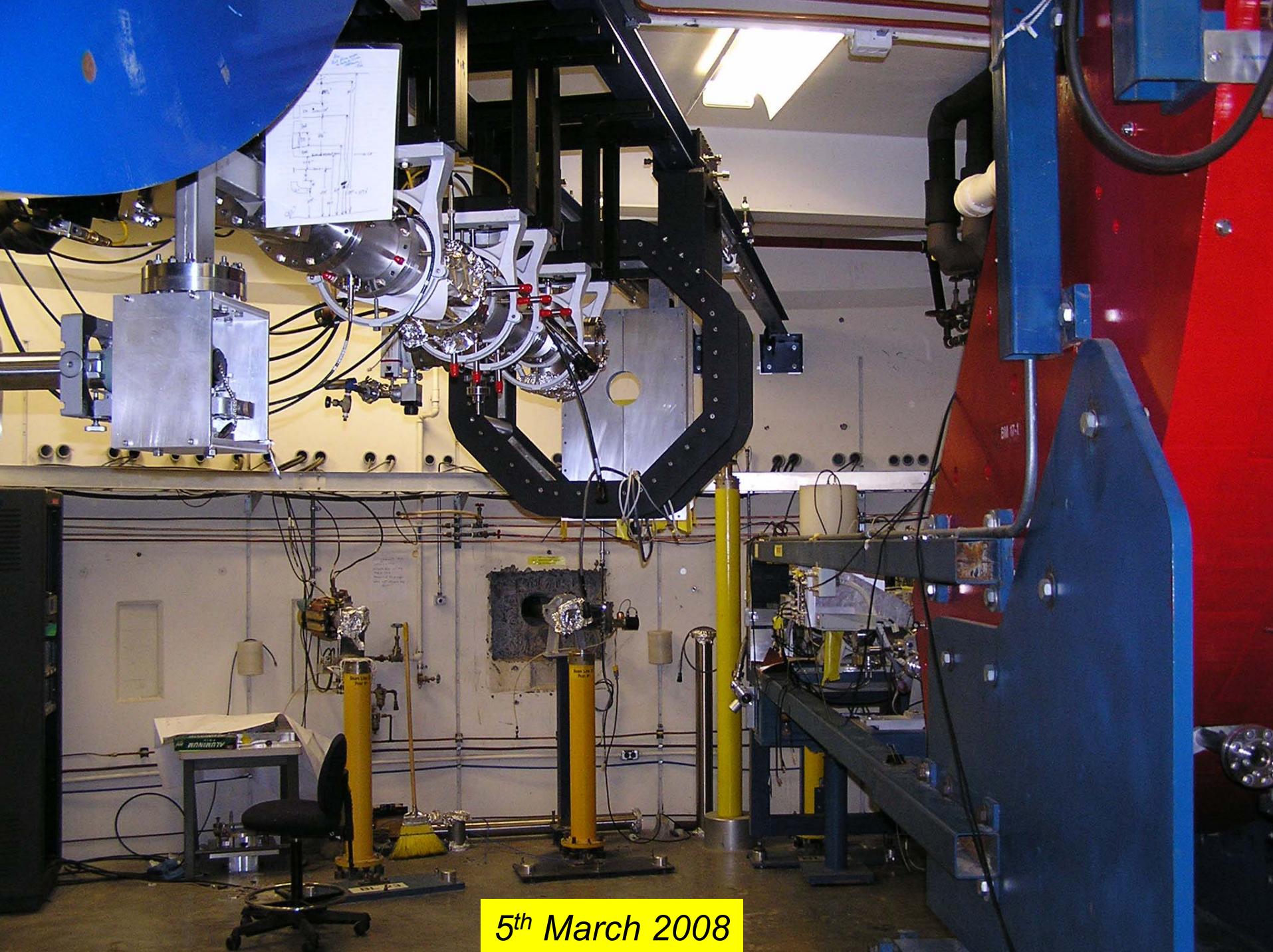
**profiting from all HRIBF beam purification methods (except “ranging-out”)**

**ultra-thin foil MCP : time correlations with implanted ions**



LeRIBSS construction should be finished in Spring 2008

New fast Moving Tape Collector and CARDS detector support from LSU !!

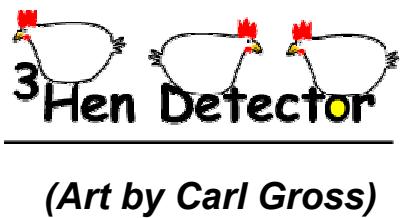


5<sup>th</sup> March 2008

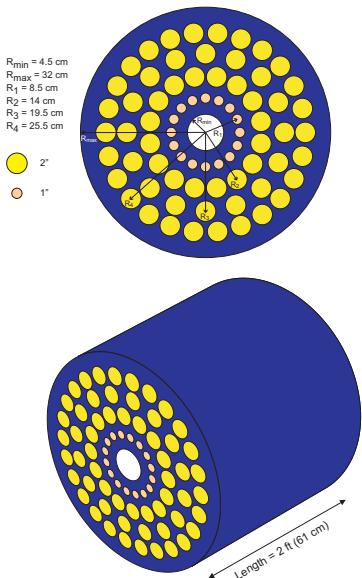
**new equipment enhancing LeRIBSS capabilities**

**nearly 80% efficient “<sup>3</sup>Hen” neutron counter**

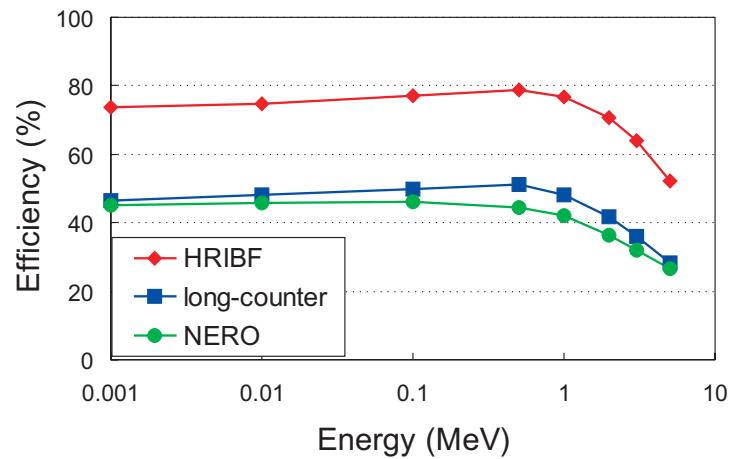
✓ purchase orders for all <sup>3</sup>He tubes /ORNL/ and MESYTEC preamps /UTK/



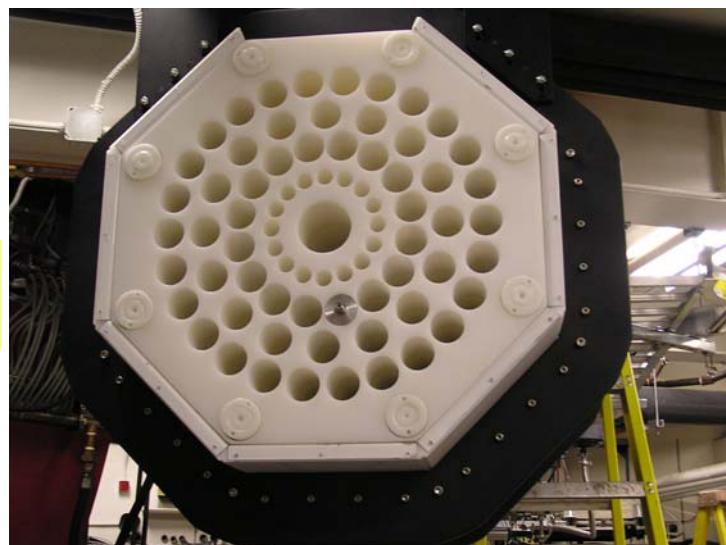
(Art by Carl Gross)



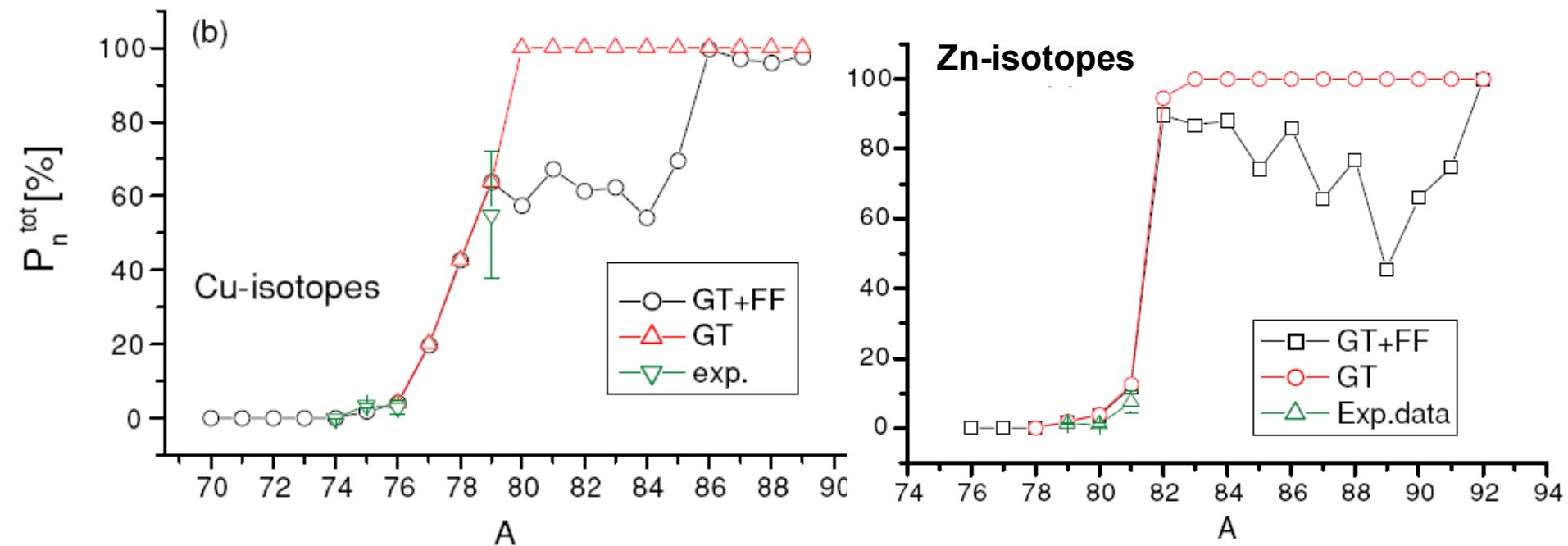
HRIBF, Long-counter, and NERO Neutron Efficiency



**UNIRIB  
LSU - Mississippi**



Efficient beta-neutron counting is critical for enhancing our “discovery potential” !!!!



from I.N. Borzov, Phys.Rev. C71, 065801, 2005

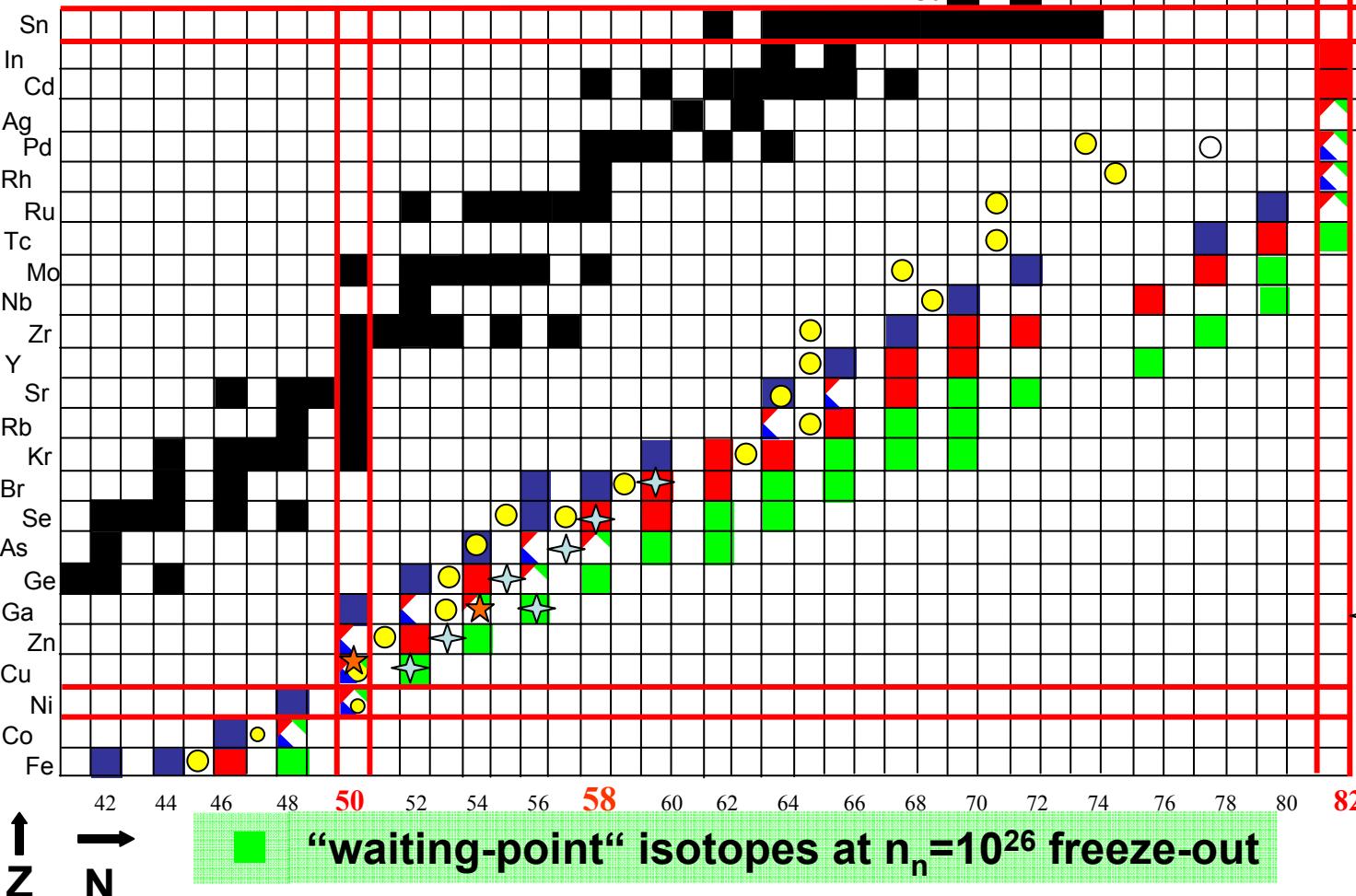
**Warning :** no  $\beta-\gamma$ , only  $\beta-n$  ( $\gamma$ ) measurements ! ????

# r-process “FK<sup>2</sup>L boulevard“ for $n_n = 10^{20}, 10^{23}, 10^{26}$ neutrons/cm<sup>3</sup>

**F.-K. Thielemann, K.-L. Kratz**, in contr. to Astrophysics Workshop, Oak Ridge 2006

- most n-rich isotope with *reported* half-life

Chart of Nuclei, 7<sup>th</sup> ed., August 2006



★ **HRIBF Nov'06**  
postaccelerated  
 $^{79}\text{Cu}$  and  $^{85}\text{Ga}$   
at 0.1-0.2 pps

★ **HRIBF LeRIBSS**  
“beyond  $^{78}\text{Ni}$  ”

# *Eksperymenty z pomocą LeRIBSS (rozпадy $\beta$ produktów rozszczepienia $^{238}\text{U}$ )*

*Nowe nuklidy procesu r :  $T_{1/2}(\beta)$  oraz  $I_{\beta n}$   
 $^{79-82}\text{Cu}$ ,  $^{80-83}\text{Zn}$ ,  $^{85-87}\text{Ga}$ ,  $^{86-87}\text{Ge}$ ,  $^{88-90}\text{As}$*

*Stany jednoczastkowe powyżej  $^{78}\text{Ni}$   
 $\nu s_{1/2}$  w  $N=51$   $^{81}\text{Zn}$  oraz  $N=53$   $^{85}\text{Ge}$  (nowa magiczna liczba  $N=58$  ?),  
poziomy  $2^+$  w  $N=52$   $^{82}\text{Zn}$  oraz  $N=54$   $^{88}\text{Se}$*

*Funkcja nasilenia  $\beta$  i rola przejsc “first-forbidden”  
 $^{87-94}\text{Br}$  isotopes (chemicznie czyste wiazki i LeRIBSS!)*

*obszar  $^{132}\text{Sn}$   
Czyste wiazki izotopow I, Ag oraz SnS*

*najbardziej egzotyczne (nowe)  
isotopy n-nadmiarowe*

*produkowane w rozszczepieniu  $^{238}\text{U}$  przez 50 MeV protony*

hribr

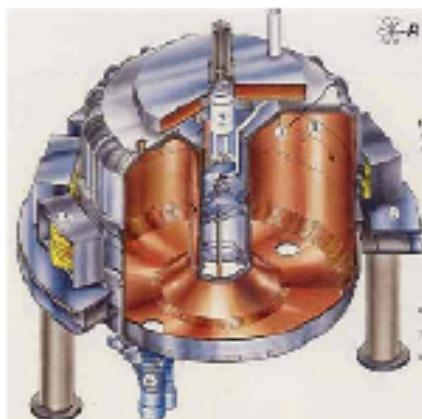
(pure)

It's the beam, stupid!

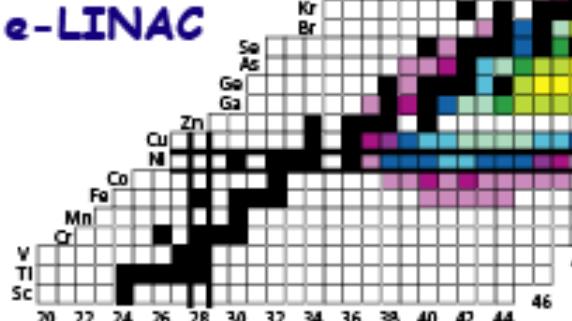
(paraphrase James Carville of the 1992 Clinton campaign)

HRIBF accelerated beam-on-target intensities  
(produced via photofission of  $^{238}\text{U}$  at  $10^{13}$  fissions/sec)

Possible e- driver:  
Rhodotron by IBA



or a 50 MeV  
e-LINAC



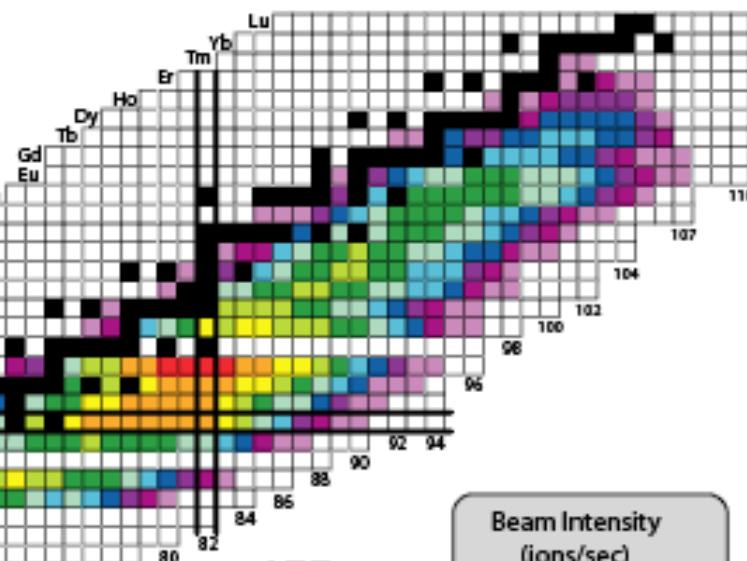
~1 ion/sec

Holifield Radioactive Ion Beam Facility

CJ Gross

Oak Ridge National Laboratory

Lu  
Yb  
Tm  
Er  
Dy  
Ho  
Gd  
Eu  
Sm  
Pm  
Nd  
Pr  
La  
Cs  
Ba  
Xe  
I  
Sn  
In  
Cd  
Ag  
Pd  
Rh  
Ru  
Tc  
Nb  
Mo  
Zr  
Y  
Sr  
Rb  
Kr  
Br  
Se  
As  
Ge  
Ga  
Zn  
Cu  
Ni  
Co  
Fe  
Mn  
Cr  
V  
Ti  
Sc  
20 22 24 26 28 30 32 34 36 38 40 42 44 46  
20 22 24 26 28 30 32 34 36 38 40 42 44 46  
48 50 52 54 56 58 60 62 64 66 68 70 72 74 76 78 80 82 84 86 88 90 92 94 96 98 100 102 104 106 108 110



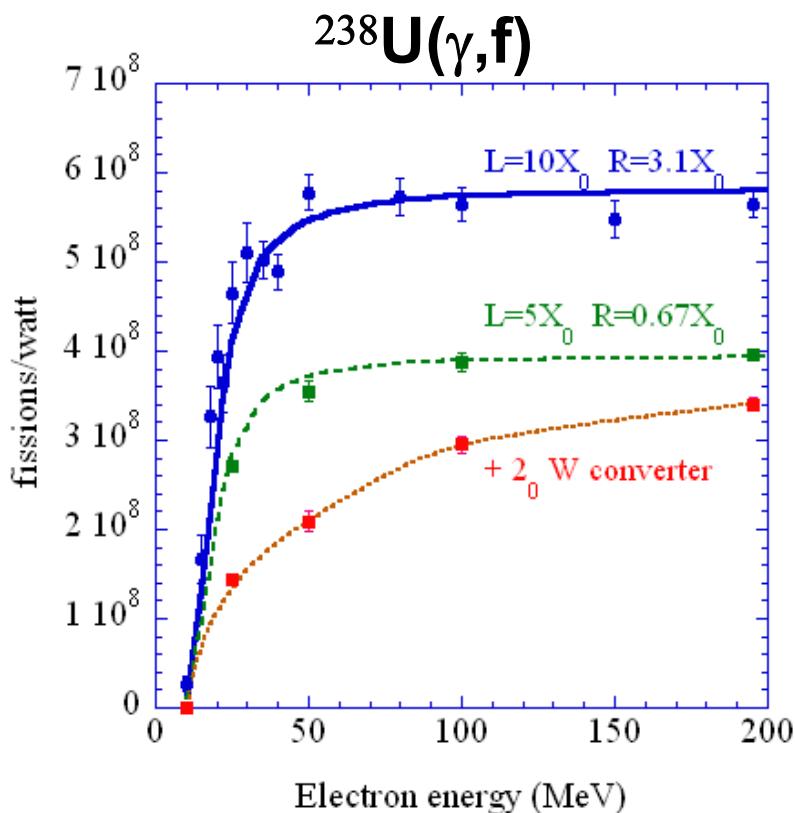
Element	Beam Intensity (ions/sec)
$^{155}\text{Ce}$	$10^9$
$^{129}\text{Ag}$	$10^8$
$^{134}\text{In}$	$10^7$
$^{138}\text{Sn}$	$10^6$
$^{92}\text{Se}$	$10^5$
$^{96}\text{Br}$	$10^4$
$^{80}\text{Cu}$	$10^3$
$^{86}\text{Ga}$	$10^2$
$^{89}\text{Ge}$	10
$^{91}\text{As}$	1
$^{102}\text{Rb}$	0.001 - 0.1
$^{145}\text{I}$	$10^9$
$^{149}\text{Cs}$	$10^8$
$^{150}\text{Ba}$	$10^7$
$^{153}\text{La}$	$10^6$
$^{142}\text{Te}$	$10^5$
$^{140}\text{Sb}$	$10^4$
$^{168}\text{Eu}$	$10^3$
$^{165}\text{Sm}$	$10^2$
$^{168}\text{Gd}$	10
$^{171}\text{Tb}$	1
$^{171}\text{Dy}$	0.001 - 0.1
$^{177}\text{Ho}$	$10^9$

Beam Intensity  
(ions/sec)



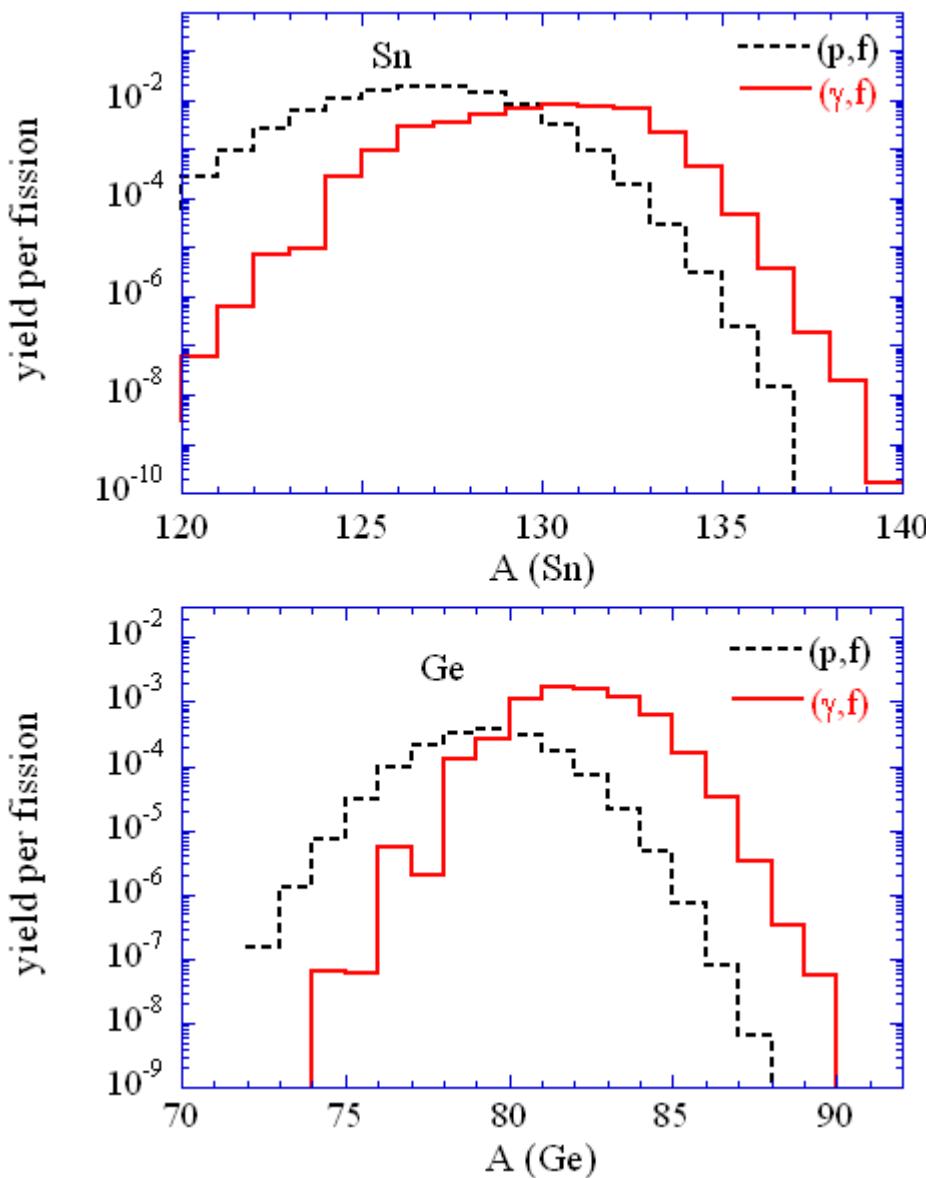
# Photofission yields

- $10^{13}$  f/s “easily” achieved
- About 20x current HRIBF
- But real gain  $\gg 20x$



(p,f) systematics from Tsukada

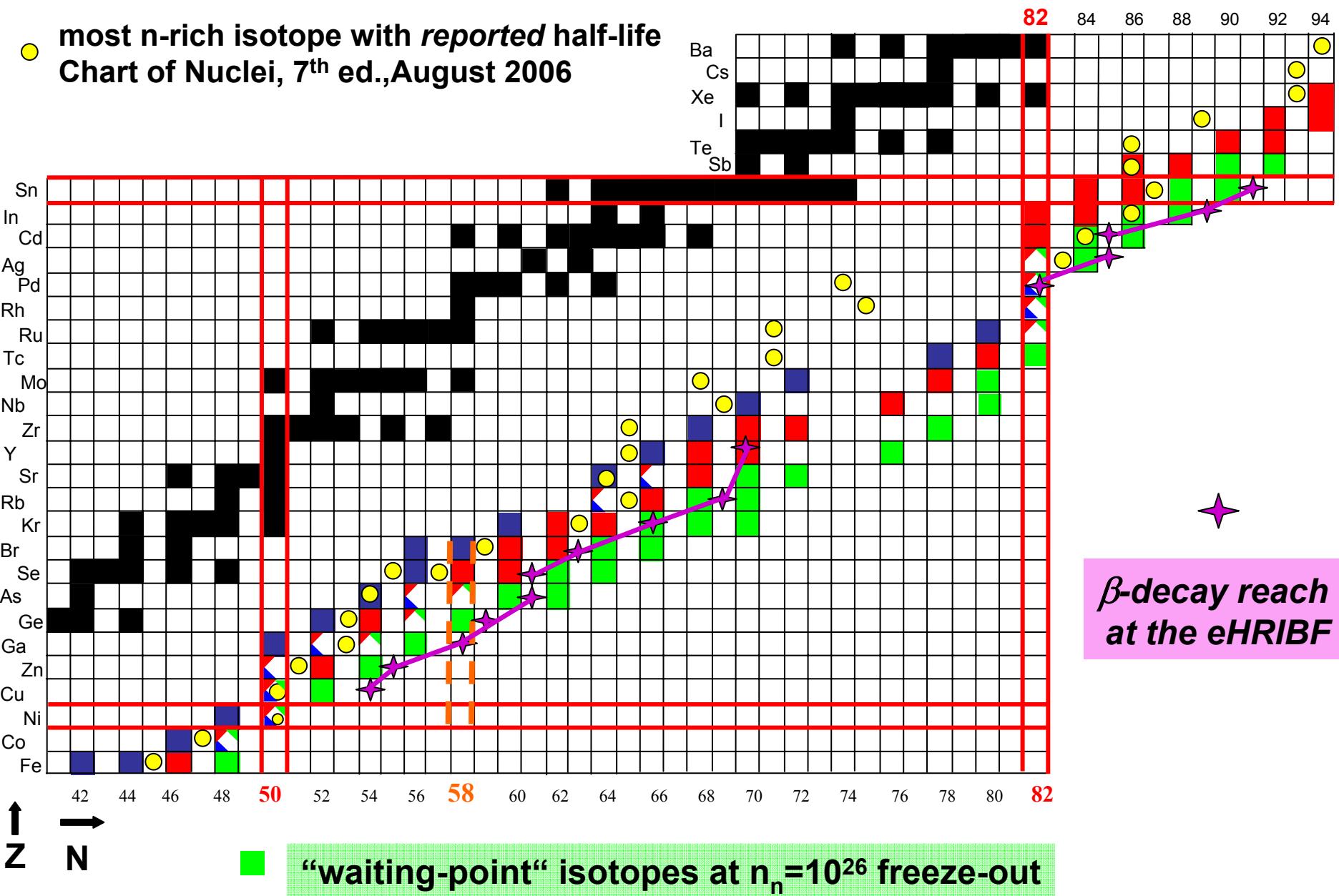
( $\gamma$ ,F) from ORNL systematics + Jyvaskyla model



# r-process for $n_n = 10^{20}, 10^{23}, 10^{26}$ neutrons/cm<sup>3</sup>

F.-K. Thielemann, K.-L. Kratz, P. Möller, et al., *AstroPhys. J.* 403, 1993; *Phys. Rev. C* 67, 2003

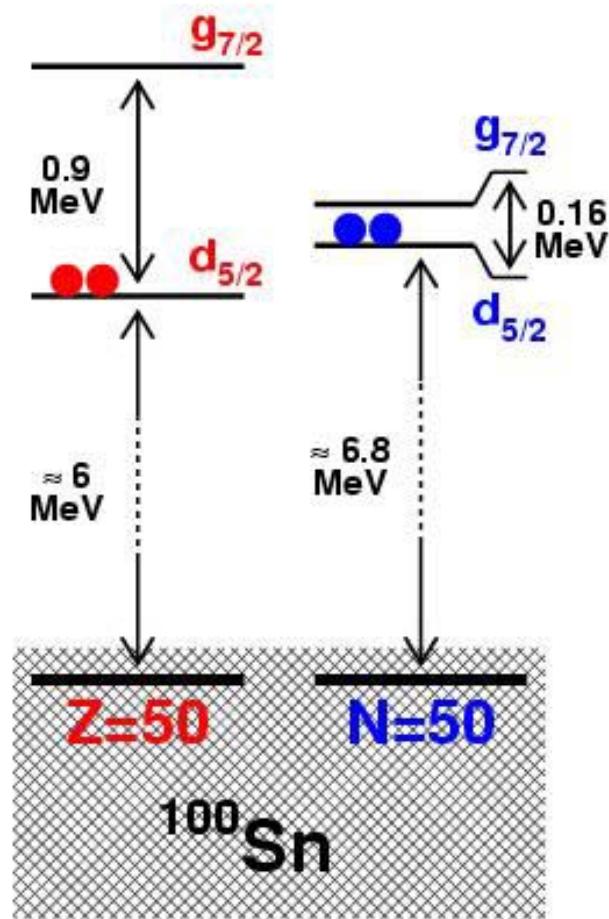
- most n-rich isotope with *reported* half-life
- Chart of Nuclei, 7<sup>th</sup> ed., August 2006



$$^{104}\text{Te} = ^{100}\text{Sn} + \alpha$$

$\alpha$  made out of  $\pi$  and  $\nu$   
on the same orbitals

*enhanced “ $\alpha$ -preformation”*

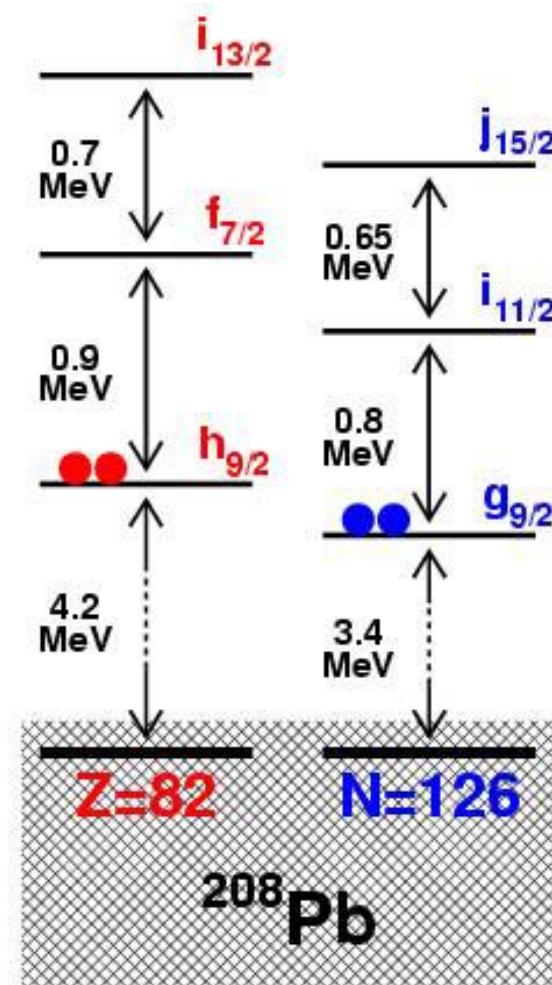


*“superallowed”  $\alpha$ -decay*

Macfarlane and Siivola, PRL 14, 1114, 1965

$$^{212}\text{Po} = ^{208}\text{Pb} + \alpha$$

$\alpha$  made out of  $\pi$  and  $\nu$   
on different orbitals



*present reference  $\alpha$ -decay*

# New alpha emitters above $^{100}\text{Sn}$

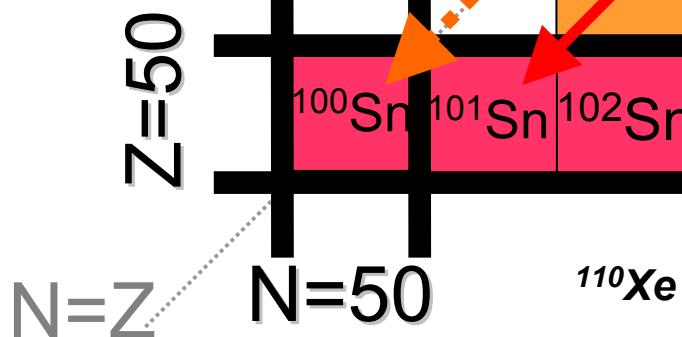
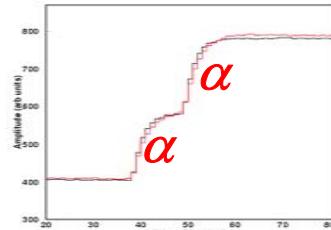
## Physics :

“superallowed”  $\alpha$ -decay ( $^{104}\text{Te} \rightarrow ^{100}\text{Sn}$  replacing  $^{212}\text{Po} \rightarrow ^{208}\text{Pb}$   $\alpha$ -reference)  
 single particle properties near doubly- magic  $^{100}\text{Sn}$   
 rp-process termination region

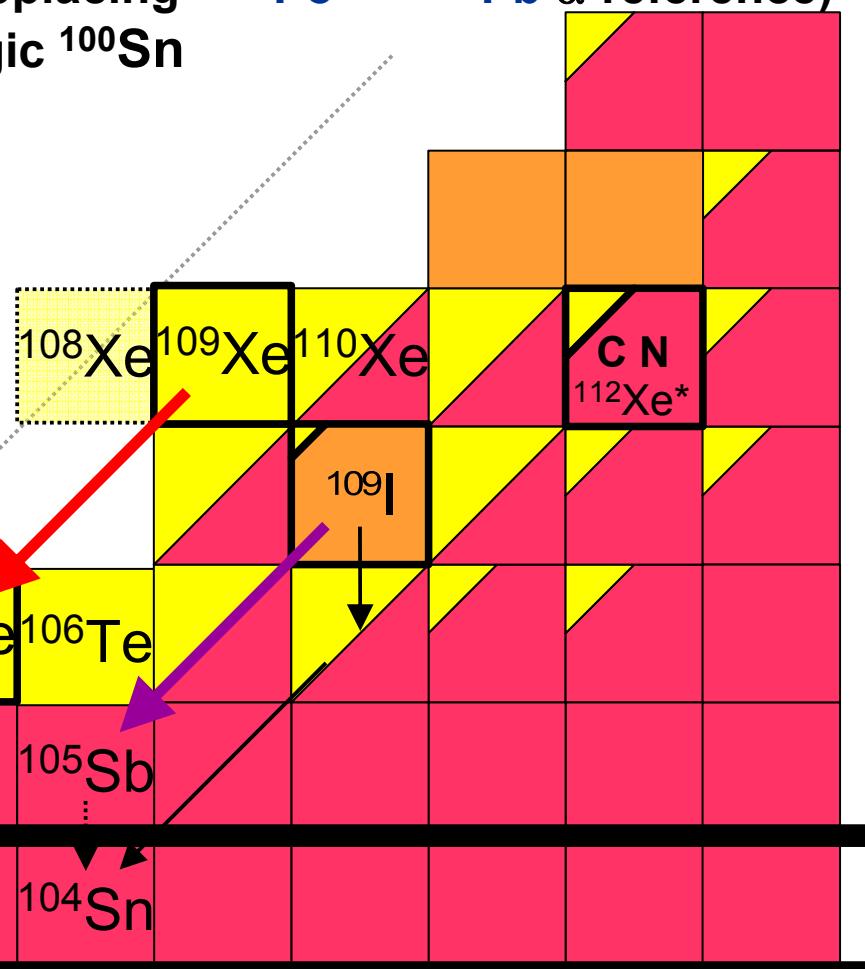


Liddick et al, PRL 97, 2006, 082501

$$T_{1/2}(^{109}\text{Xe})/T_{1/2}(^{105}\text{Te}) = 20\,000 : 1$$



$^{110}\text{Xe} \rightarrow ^{106}\text{Te} \rightarrow ^{102}\text{Sn}$ , Schardt et al., NP A368, 1981

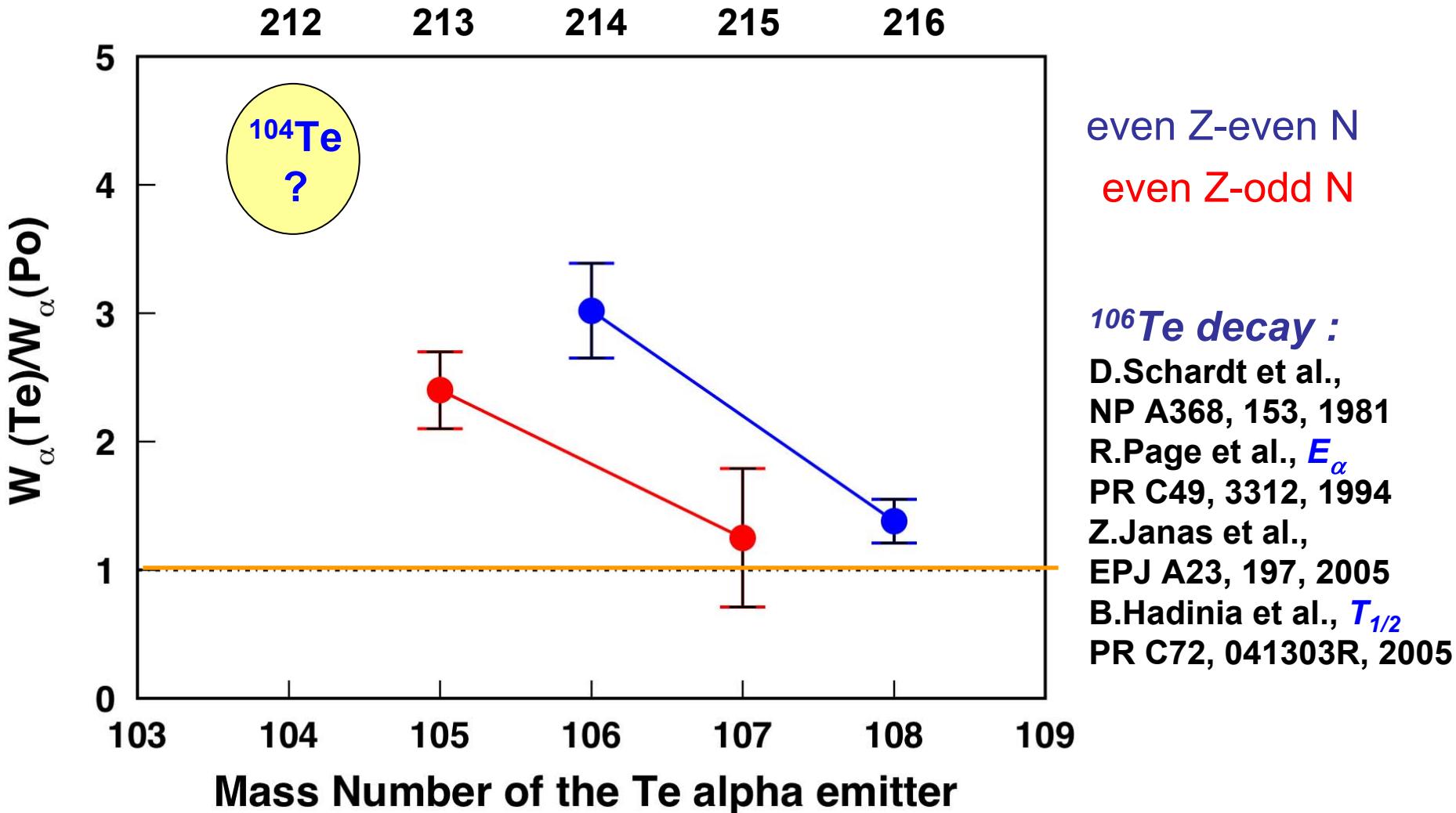


# Comparison of the reduced decay width $W_\alpha$

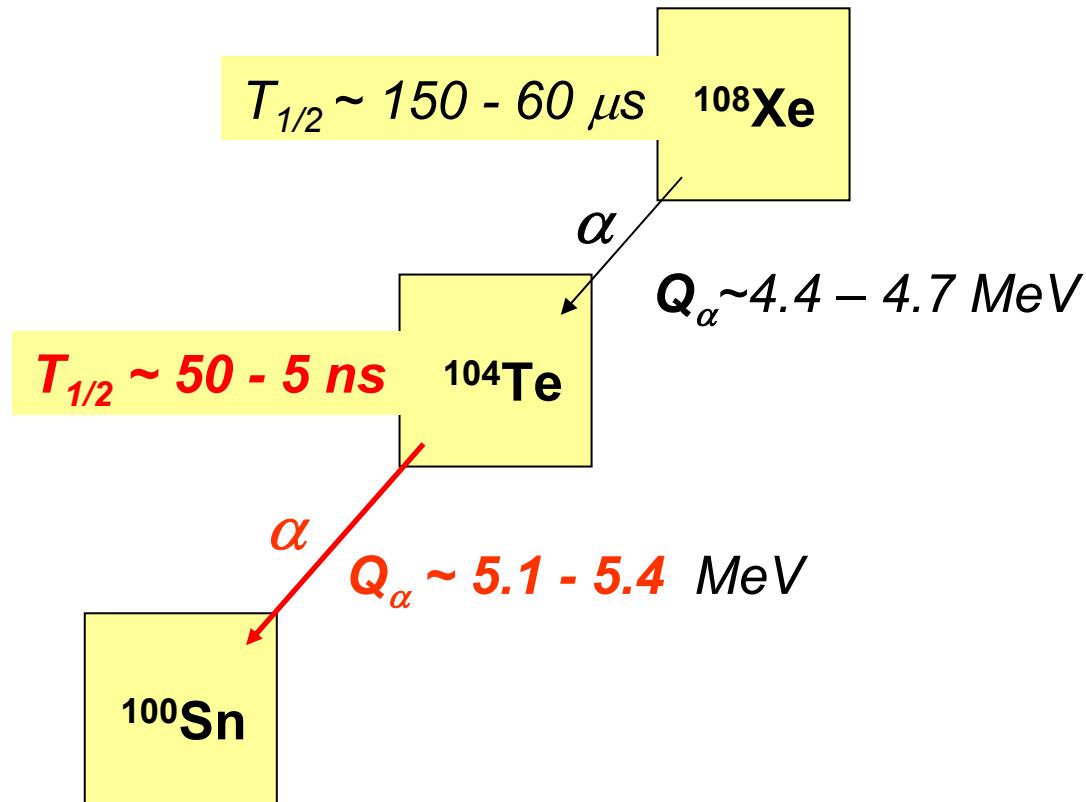


(see also : Xu,Ren,PR C74,2006, 037302 and P.Mohr, EPJ A31,2007,23)

Mass Number A of the Po alpha emitter



Poszukiwanie rozpadów alfa  $^{108}\text{Xe} \rightarrow ^{104}\text{Te} \rightarrow ^{100}\text{Sn}$  /R.Grzywacz, R.Page/

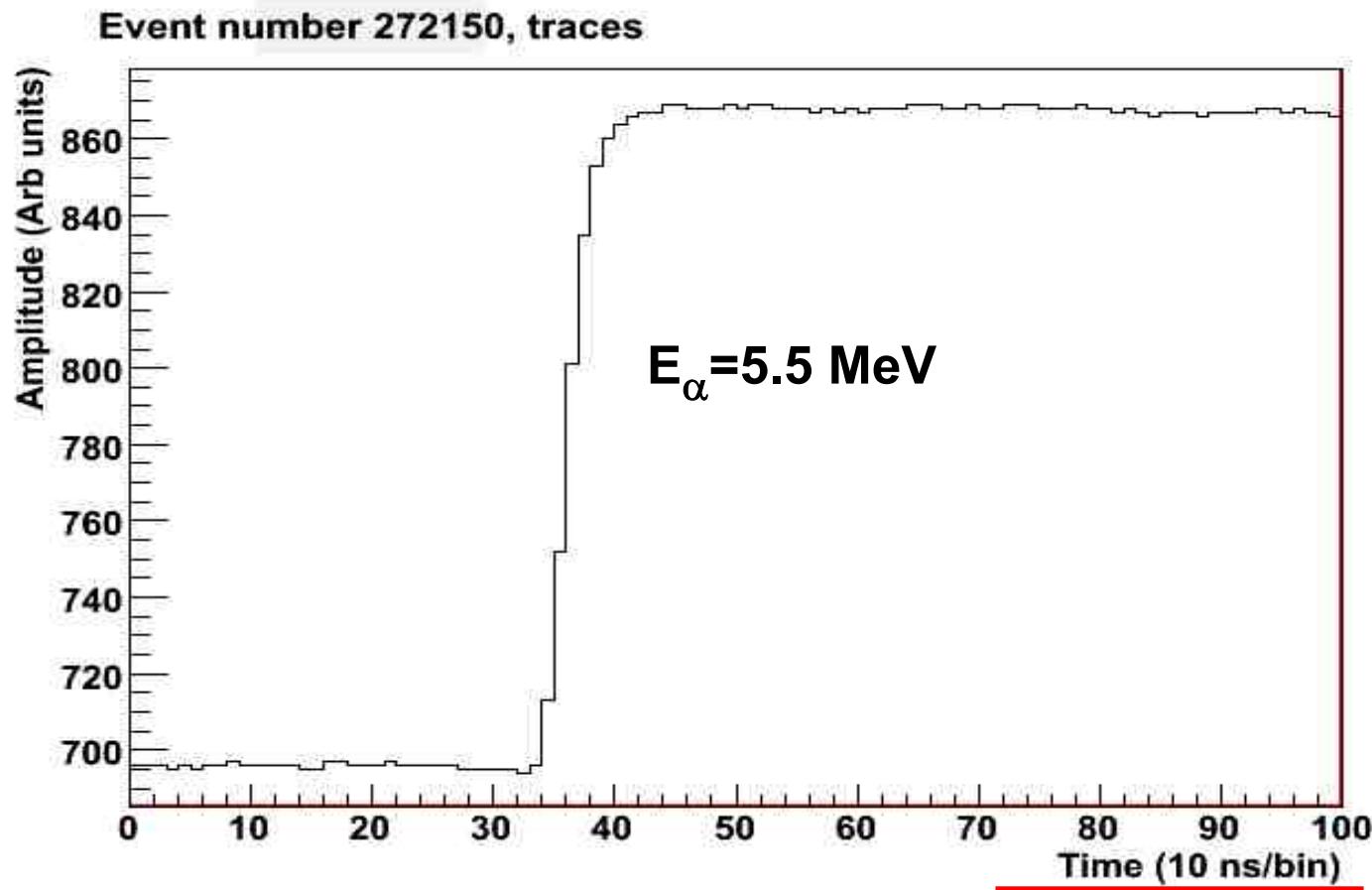


A.Korgul et al., PR C77, 034301, 2008:

100 godzin (HRIBF)  $\rightarrow$  20 jonów  $^{108}\text{Xe}$  (sygnały  $E_{\alpha 1} + E_{\alpha 2} = 9 - 10 \text{ MeV}$ )

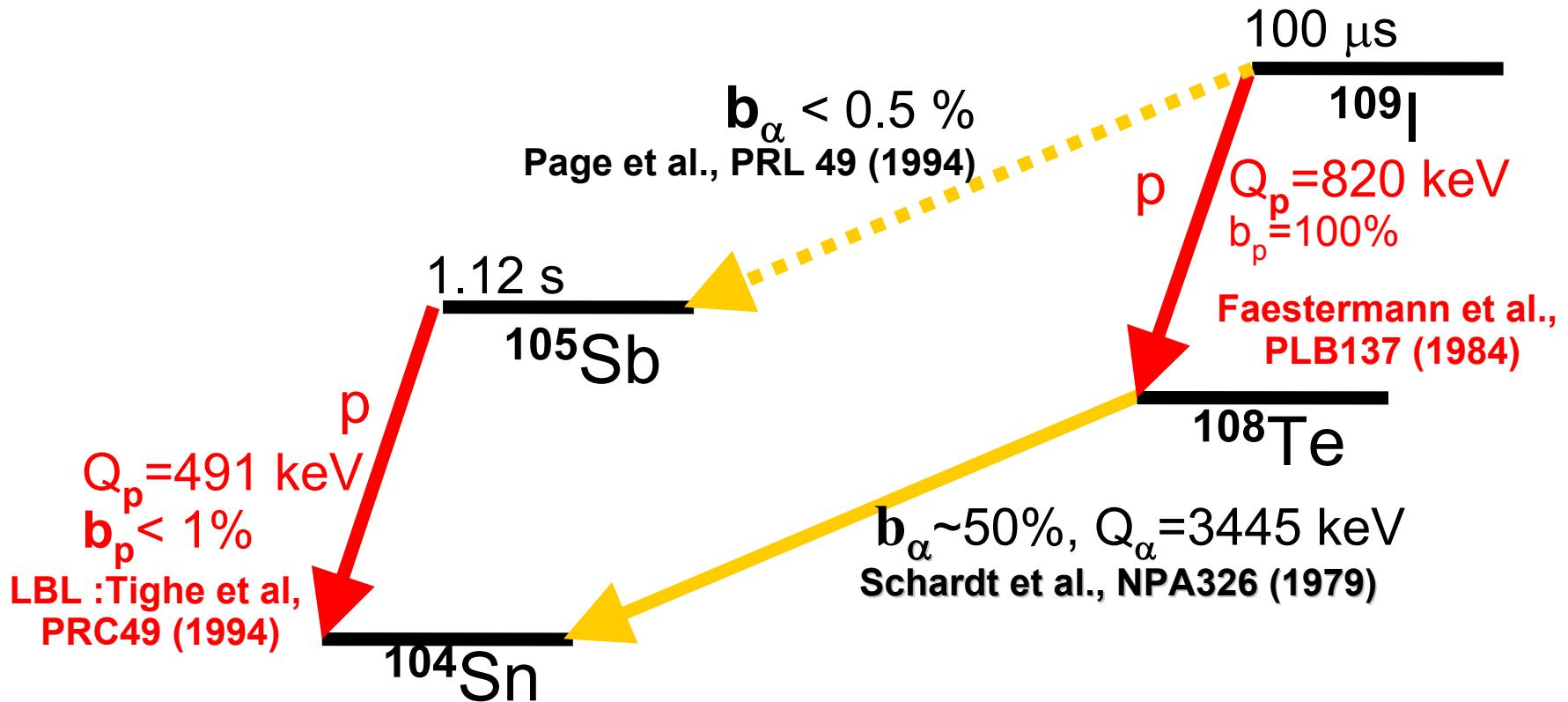
- wiązka  $^{58}\text{Ni}$ , 50 pnA, 240 MeV (ORNL Tandem) ✓
- tarcza  $^{54}\text{Fe}$  (J. Szerypo, Monachium, kolaboracja UNIRIB) ✓
- szybkie przedwzmacniacze ~ 20 ns (R.Schneider, MESYTEC, R. Grzywacz) ✓
- elektronika 100 MHz  $\rightarrow$  10ns/kanał (XIA, R. Grzywacz et al., ORNL-UTK) ✓

Sean Liddick (UTK), marzec 2008 :  
100 MHz Pixie-16 oraz "20 ns" przedwzmacniacze MESYTEC



uruchomienie i sprawdzenie nowego oprogramowanie elektroniki Pixie-16  
"decay signal selector" ([XIA](#), R.Grzywacz et al.)

# Before *h r i b f* experiment:



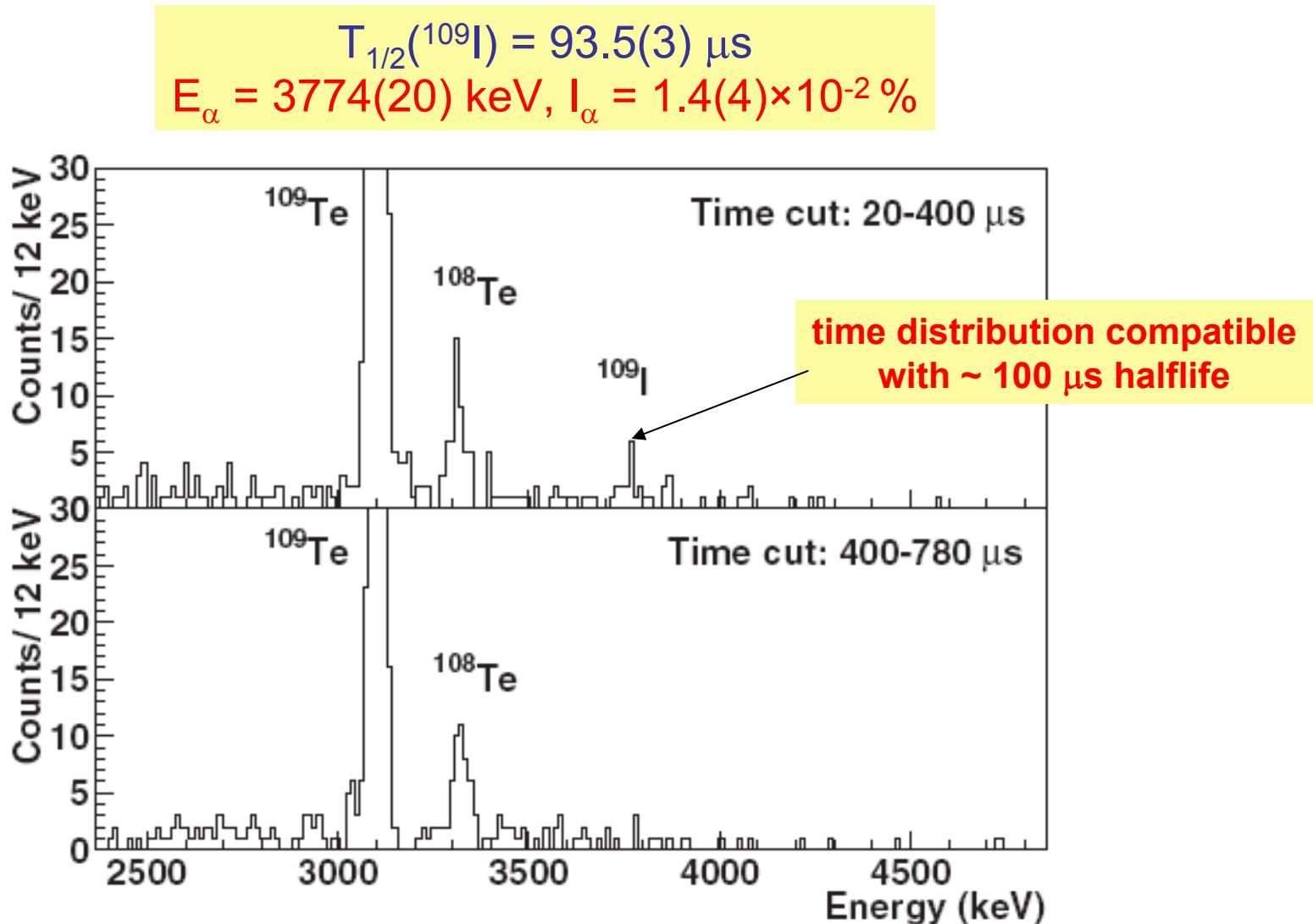
many attempts were made to detect proton emission from  $^{105}\text{Sb}$  (verify LBL result) :

Gillitzer et al., ZPA326 (1987), G. Berthes et al., GSI-87-12 (1987), J.Friese Hirscheegg 1996,

Liu et al., PRC72 (2005)

over 150 000 of  $^{109}\text{I}$  ions implanted into DSSD

directly from  $^{109}\text{I}$  protons (113 000) and from a daughter activity  $^{108}\text{Te}$   $\alpha$ -decay (70 000)



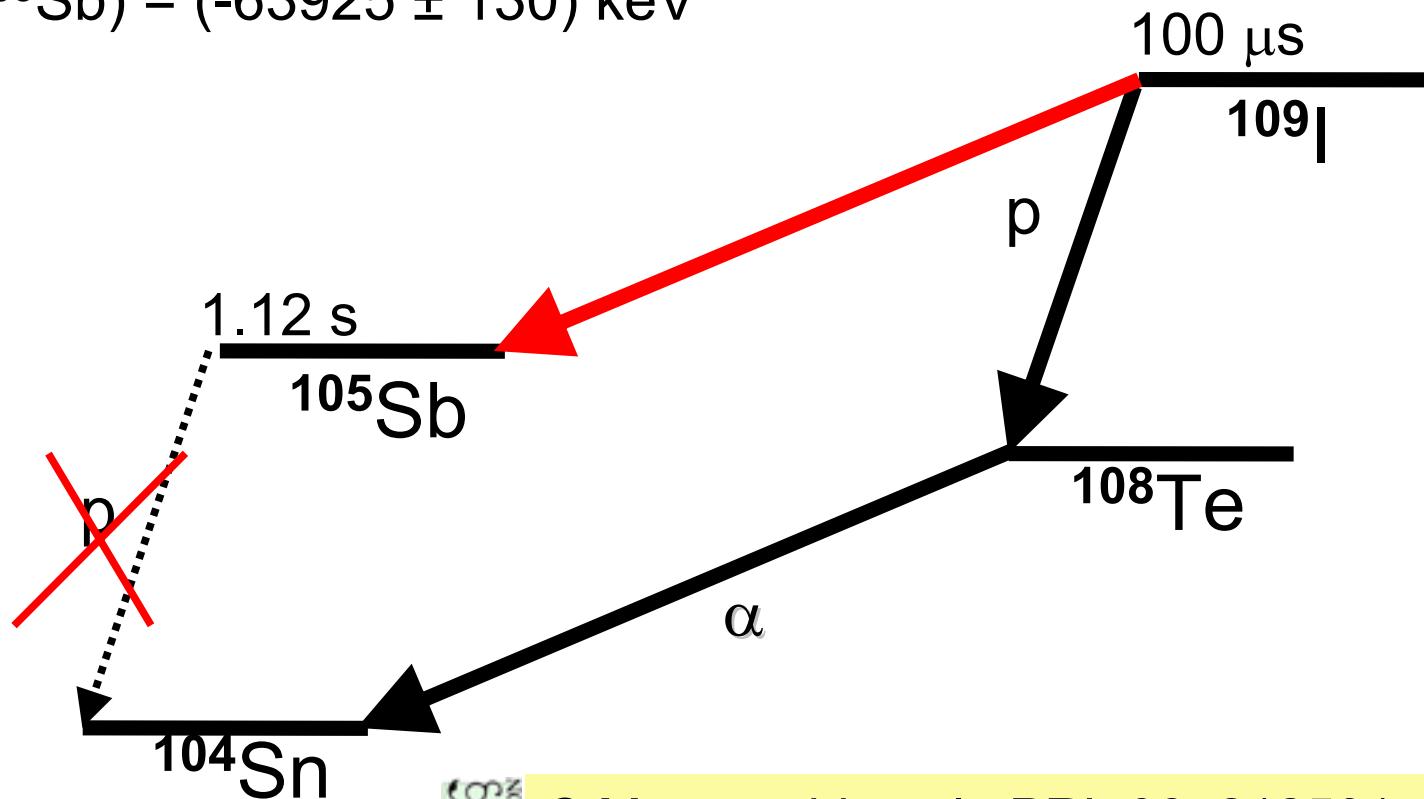
*Result : no observable proton emission from  $^{105}\text{Sb}$  !*

$$Q_\alpha(^{109}\text{I}) = 3918 \pm 21 \text{ keV}$$

$$\Rightarrow Q_p(^{105}\text{Sb}) = 356 \pm 22 \text{ keV} \quad (\text{was } 491(15) \text{ keV, Tighe et al. 1994})$$

$$\Rightarrow T_{1/2}^p \sim 10^7 \text{ s} (=116 \text{ days}) \rightarrow b_p \sim 10^{-7}$$

$$\Delta M (^{105}\text{Sb}) = (-63925 \pm 130) \text{ keV}$$



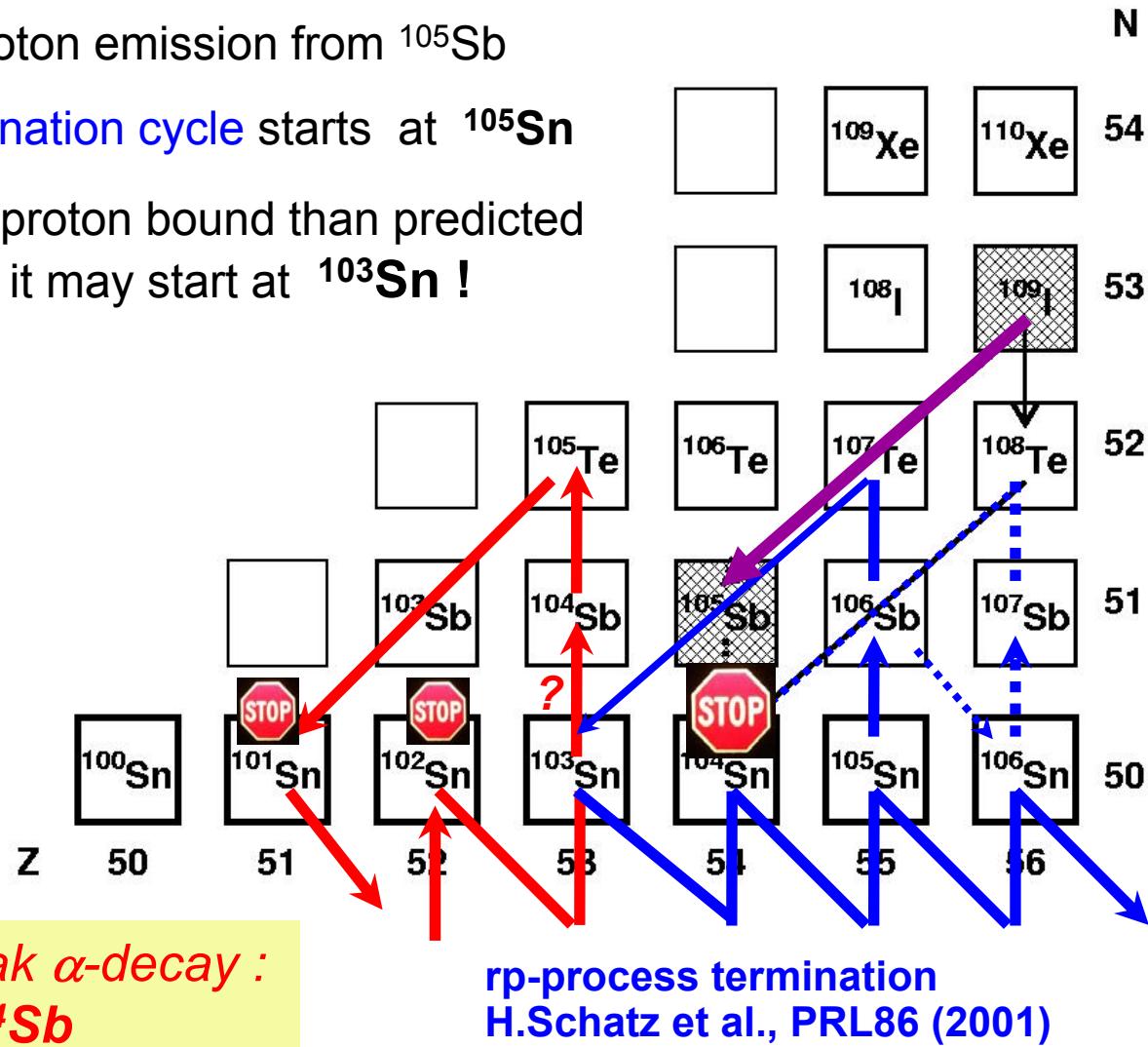
# Astrophysical relevance :



C.Mazzocchi, ..., H.Schatz,...PRL 98,212501, 2007

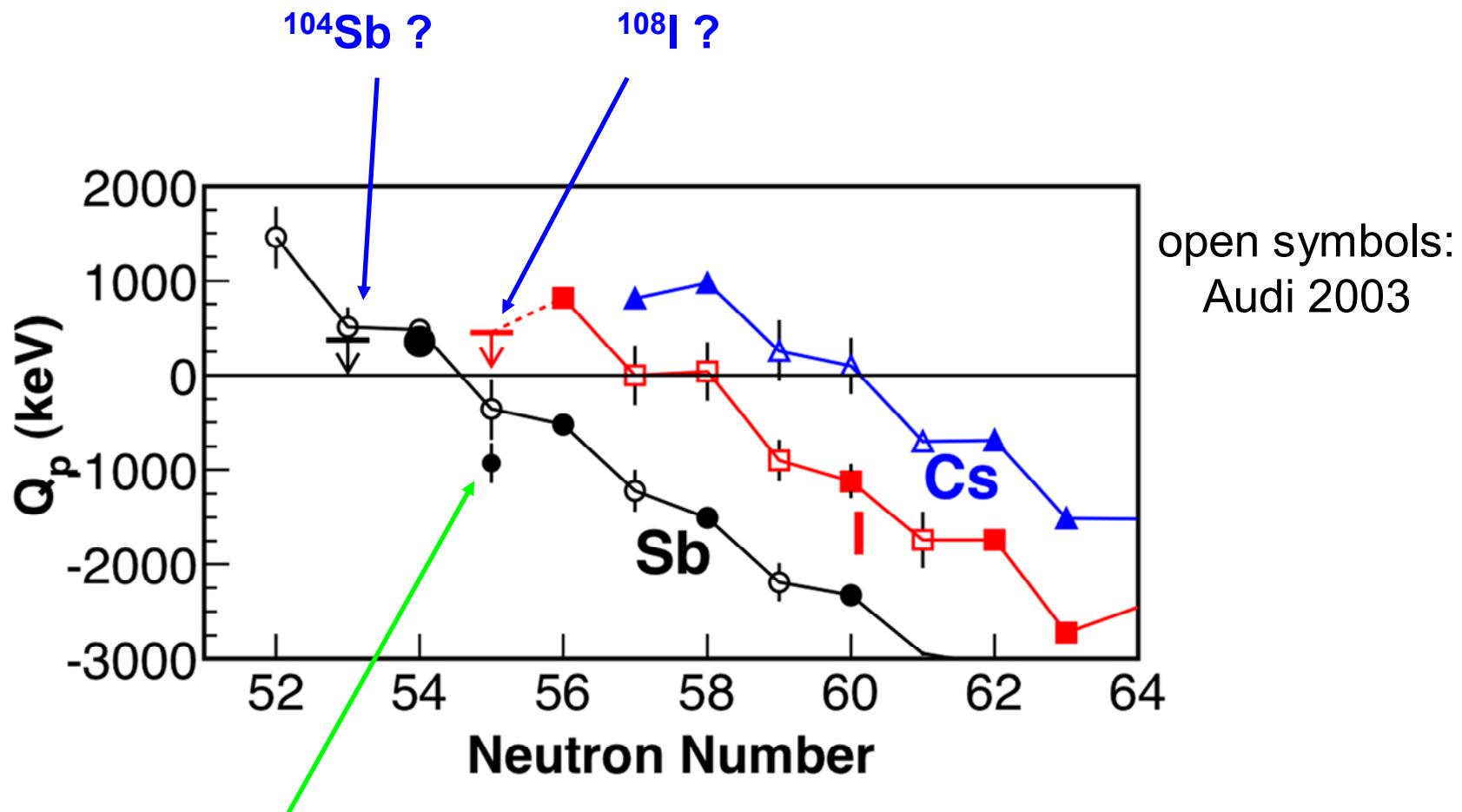


- No observable proton emission from  $^{105}\text{Sb}$
- The rp-process termination cycle starts at  $^{105}\text{Sn}$
- If  $^{104}\text{Sb}$  is much more proton bound than predicted (strong odd-even effect) it may start at  $^{103}\text{Sn}$  !



## Odd-even effect in proton decay energies $Q_p$

$Z=55 \text{ Cs}$ ,  $Z=53 \text{ I}$ ,  $Z=51 \text{ Sb}$



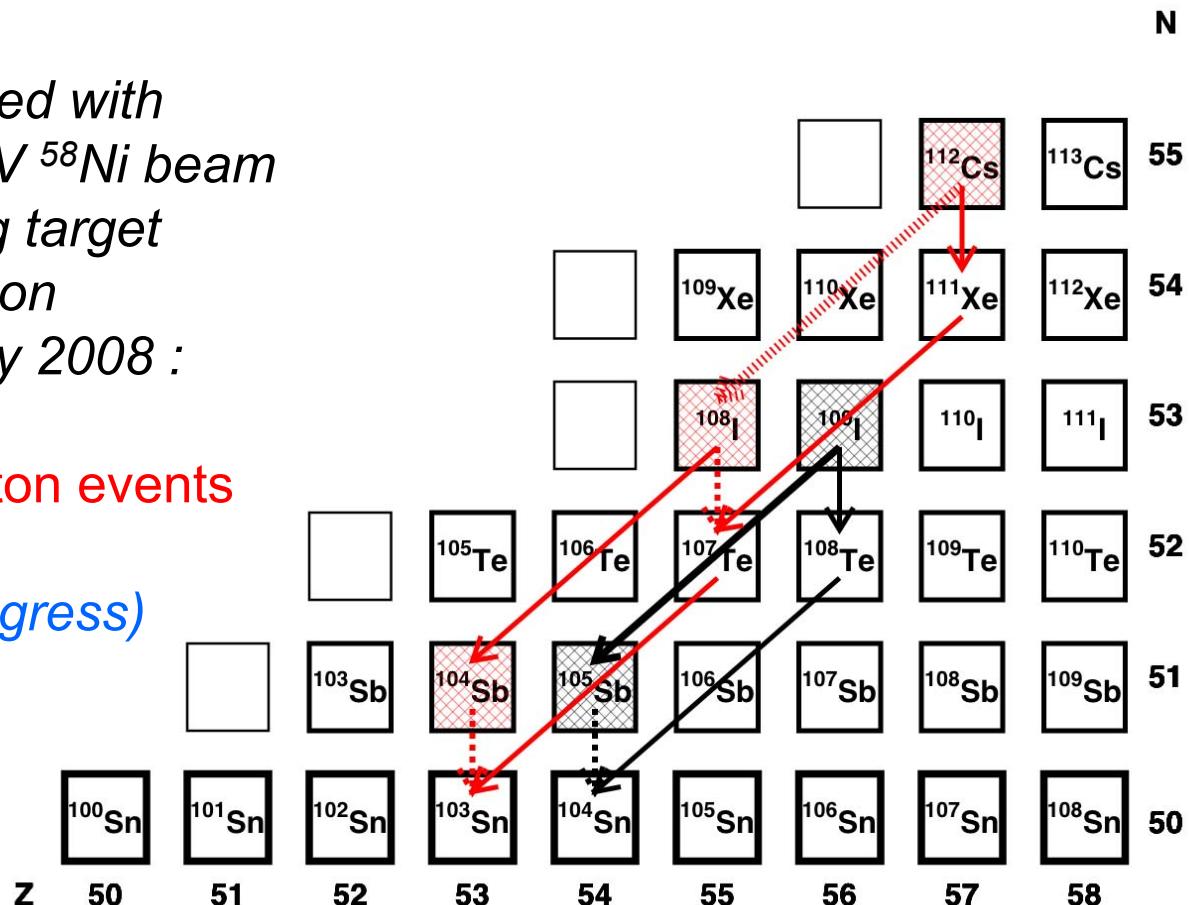
$^{106}\text{Sb}????$  Płochocki et al., Phys. Lett. 106B, 285 (1981)

$^{106}\text{Sb}$  – interesting candidate for trap exp ???

search for  $^{112}\text{Cs}$  (very weak)  $\alpha$ -decay :  
 $S_P$  of  $^{108}\text{I}$  and  $^{104}\text{Sb}$

statistics achieved with  
over 50 pnA 250 MeV  $^{58}\text{Ni}$  beam  
and  $^{58}\text{Ni}$  rotating target  
(p3n) reaction  
hrift February 2008 :

$\sim 2000$   $^{112}\text{Cs}$  proton events  
(analysis in progress)



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